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PART B
SOLAR - GEOPHYSICAL DATA

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U. S. DEPARTMENT OF COMMERCE NATIONAL BUREAU OF STANDARDS CENTRAL RADIO PROPAGATION LABORATORY BOULDER, COLORADO



### SOLAR - GEOPHYSICAL DATA

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### SOLAR - GEOPHYSICAL DATA

### INTRODUCTION

This monthly report series is intended to keep research workers abreast of the major particulars of solar activity and the associated ionospheric, radio propagation and other geophysical effects. It is made possible through the cooperation of many observatories, laboratories and agencies as recorded in the detailed description of the tables and graphs which follows. The report is prepared in the Sun-Earth Relationships Section, edited by Miss J. V. Lincoln and Mr. Dale B. Bucknam.

### I DAILY SOLAR INDICES

Relative Sunspot Numbers -- The table includes (1) the daily American relative sunspot numbers,  $R_A$ , as compiled by the Solar Division of the American Association of Variable Star Observers, and (2) the provisional daily Zürich relative sunspot numbers,  $R_Z$ , as communicated by the Swiss Federal Observatory. Because of the time required to collect and reduce the observations,  $R_A$  will normally appear one month later than  $R_Z$ .

The relative sunspot number is an index of the activity of the entire visible disk. It is determined each day without reference to preceding days. Each isolated cluster of sunspots is termed a sunspot group and it may consist of one or a large number of distinct spots whose size can range from 10 or more square degrees of the solar surface down to the limit of resolution (e.g. 1/8 square degrees). The relative sunspot number is defined as R=K(10g+s), where g is the number of sunspot groups and s is the total number of distinct spots. The scale factor K (usually less than unity) depends on the observer and is intended to effect the conversion to the scale originated by Wolf. The observations for sunspot numbers are made by a rather small group of extraordinarily faithful observers, many of them amateurs, each with many years of experience. The counts are made visually with small, suitably protected telescopes.

Final values of  $R_Z$  appear in the IAU <u>Quarterly Bulletin on Solar Activity</u>, the <u>Journal of Geophysical Research</u> and elsewhere. They usually differ slightly from the provisional values. The American numbers,  $R_\Delta$ , are not revised.

Solar Flux Values, 2800 Mc -- The table also lists the daily values of solar flux at 2800 Mc recorded in watts/M<sup>2</sup>/cycle/second bandwidth (x 10<sup>-22</sup>) in two polarizations by the National Research Council at Ottawa, Canada. These solar radio noise indices are being published in accordance with CCIR Report 25 that a basic solar index for ionospheric propagation should be measured objectively and "preferably refer to a property of the sun such as radiation flux which has direct physical relationship to the ionosphere."

Graph of Sunspot Cycle -- The graph illustrates the recent trend of Cycle 19 of the 11-year sunspot cycle and some predictions of the future level of activity. The customary "12-month" smoothed index, R, is used throughout, the data being final  $R_Z$  numbers except for the current year. Predictions shown are those made for one year after the latest available datum by the method of A. G. McNish and J. V. Lincoln (Trans. Am. Geophys. Union, 30, 673-685, 1949) modified by the use of regression coefficients and mean cycle values recomputed for Cycles 8 through 18. Cycle 19 began April 1954, when the minimum  $\overline{R}$  of 3.4 was reached.

### II SOLAR CENTERS OF ACTIVITY

<u>Calcium Plage and Sunspot Regions</u> -- The table gives particulars of the centers of activity visible on the solar disk during the preceding month. These are based on estimates made and reported on the day of observation and are therefore of limited reliability.

The table gives the heliographic coordinates of each center (taken as the calcium plage unless two or more significantly and individually active sunspot groups are included in an extended plage) in terms of the Greenwich date of passage of the sun's central meridian (CMP) and the latitude; the serial number of the plage as assigned by McMath-Hulbert Observatory; the serial number of the center in the previous solar rotation, if it is a persisting region; particulars of the plage at CMP: area, central intensity; a summary of the development of the plage during the current transit of the disk, where b = born on disk,  $\ell$  = passed to or from invisible hemisphere, d = died on disk, and  $\prime$  = increasing, - = stable, \ = decreasing; and age in solar rotations; particulars of the associated sunspot group, if any, at CMP: area and spot count and the summary of development during the current disk transit, similar to the above. The unit of area is a millionth of the area of a solar hemisphere; the central intensity of calcium plages is roughly estimated on a scale of l = faint to 5 = very bright.

Calcium plage data are available through the cooperation of the McMath-Hulbert Observatory of the University of Michigan and the Mt. Wilson Observatory. The sunspot data are compiled from reports from the U.S. Naval Observatory, Mt. Wilson Observatory, and from reports from Europe and Japan received through the daily Ursigram messages.

Coronal Line Emission Indices -- In the table are summarized solar coronal emission intensity indices for the green (Fe XIV at  $\lambda5303$ ) and red (Fe X at  $\lambda6374$ ) coronal lines. The indices are based on measurements made at 5° intervals around the periphery of the solar disk by the High Altitude Observatory at Climax, Colorado, and by Harvard University observers at Sacramento Peak (The USAF Upper Air Research Observatory at Sunspot, New Mexico, under contract AF 19(604)-146). The measurements are expressed as the number of millionths of

an Angstrom of the continum of the center of the solar disk (at the same wavelength as the line) that would contain the same energy as the observed coronal line. The indices have the following meanings:

 $G_6$  = mean of six highest line intensities in quadrant for  $\lambda 5303$ .

 $R_6 = \text{same for } \lambda 6374.$ 

 $G_1$  = highest value of intensity in quadrant, for  $\lambda 5303$ .

 $R_1 = \text{same for } \lambda 6374.$ 

The dates given in the table correspond to the approximate time of CMP of the longitude zone represented by the indices. The actual observations were made for the North East and South East quadrants 7 days before; for the South West and North West quadrants 7 days after the CMP date given.

To obtain rough measures of the integrated emission of the entire solar disk in either of the lines, assuming the coronal changes to be small in a half solar rotation, it is satisfactory to perform the following type of summation given in example for 15 October:

where N is the number of indices entering the summation.

Such integrated disk indices as well as integrated wholesun indices are computed for each day and are published quarterly in the "Solar Activity Summary" issued by the High Altitude Observatory at Boulder, Colorado. In the same reports are given maps of the intensity distribution of coronal emission derived from all available Climax and Sacramento Peak observations, as well as other information on solar activity, such as maps made from daily limb prominence surveys in  ${\rm H}\alpha$  and notes regarding the history of active regions on the solar disk.

Preliminary summaries of solar activity, prepared on a fast schedule, are issued Friday of each week from High Altitude Observatory in conjunction with CRPL and include solar activity through the preceding day. These are useful to groups needing information on the current status of activity on the visible solar disk, but are not recommended for research uses unless such a prompt schedule of reporting is essential. The same information is included in the subsequent quarterly reports, with extensive additions, corrections and evaluations.

Optical Observations -- The table presents the preliminary record of solar flares as reported to the CRPL on a rapid schedule at the sacrifice of detailed accuracy. Definitive and complete data are published later in the Quarterly Bulletin on Solar Activity, I.A.U., in various observatory publications and elsewhere. The present listing serves to identify and roughly describe the phenomena observed.

Reporting directly to the CRPL are the following observatories: McMath-Hulbert, Wendelstein, Sacramento Peak, Mitaka and Swedish Astrophysical Station on Capri. The remainder report through the URS Igram centers or are available through the IGY World Data Center for Solar Activity in Boulder. Observations are in the light of the center of the H-alpha line unless noted otherwise. The reports from Sacramento Peak, New Mexico (communicated to CRPL by the High Altitude Observatory at Boulder) are from observations at the USAF Upper Air Research Observatory at Sunspot, New Mexico, by Harvard University observers, under contract AF 19(604)-146.

For each flare are listed the reporting observatory, the date, beginning and ending times, time of maximum phase, the heliographic coordinates in degrees, McMath serial number of the region, duration, the flare importance on the IAU scale of 1- to 3+, observing conditions where 1 means poor, 2 fair and 3 good, time of measurement for tabulated width of Ha or tabulated area, measured (i.e. projected) maximum area in square degrees, corrected maximum area in square degrees which equals measured area times secant h where h is the heliocentric angle, maximum effective line-width in Ha expressed in Angstroms, and maximum intensity of Ha expressed in per cent of the continuous spectrum. The following symbols are used in the table:

D = Greater than F = Approximately E = Less than G = Plus

A final column lists provisionally the occurrence of simultaneous ionospheric effects as observed on selected field-strength recordings of distant high-frequency radio transmissions; a more nearly definitive list of these ionospheric effects, including particulars, appears in these reports after the lapse of a month (see below). All times are Universal Time (UT or GCT). Subflares (importance 1-) are listed by date, time of beginning and their heliographic coordinates. A graph presents intervals for which there were no patrols for flare observations from the observatories whose complete data are published in the table.

Ionospheric Effects -- SID (and GID--gradual ionospheric disturbances) may be detected in a number of ways: short wave fadeouts, enhancement of low frequency atmospherics, increases in cosmic absorption, and so forth. The table lists events that have been recognized on field-strength recordings of distant high-frequency radio transmissions.

Under a coordinated program, the staffs at the following ionospheric sounding stations contribute reports that are screened and synthesized at CRPL-Boulder: Puerto Rico, Ft. Belvoir, Va., and Anchorage, Alaska (CRPL Stations: PR, BE, AN): Huancayo, Peru (CRPL-Associated Laboratory: HU); and White Sands, N. Mex., Adak, Alaska, and Okinawa (U.S. Signal Corps Stations: WS, AD, OK). McMath-Hulbert Observatory (MC) also contributes such reports. In addition, reports are volunteered by RCA Communications Inc., Marconi Wireless, Netherlands Postal and Telecommunications Services, Swedish Telecommunications, and others; these usually specify times of SJD and the radio paths involved.

In the coordinated program, the abnormal fades of field strength not obviously ascribable to other causes, are described as short wave fadeouts with the following further classification:

S-SWF: sudden drop-out and gradual recovery Slow S-SWF: drop-out taking 5 to 15 minutes and

gradual recovery

G-SWF: gradual disturbance; fade irregular in either drop-out or recovery or both.

When there is agreement among the various reporting stations on the time (UT) of an event, it is accepted as a widespread phenomenon and listed in the table.

The degree of confidence in identifying the event, a subjective estimate, is reported by the stations and this is summarized in an index of certainty that the event is widespread, ranging from 1 (possible) to 5 (definite). The times given in the table for the event are from the report of a station (underlined in table) that identified it with high confidence. The criteria for the subjective importance rating assigned by each station on a scale of 1- to 3+ include amplitude of the fade, duration and confidence; greater consideration is given to reports on paths near the subsolar point in arriving at the summary importance rating given in the table.

Note: The tables of SID observed at Washington included in CRPL F-

Note: The tables of SID observed at Washington included in CRPL Freports prior to F-135 were restricted to events classed here as S-SWF.

### IV SOLAR RADIO WAVES

### 2800 Mc Observations

The data on solar radio wave events made in Ottawa, Canada by the Radio and Electrical Engineering Division of the National Research Council (A. E. Covington) at 2800 Mc (10-cm emission) are presented. Near local noon (about 1700 UT) the sensitivity of the radiometer is determined and a mean flux for the whole day calculated. These values are given in a tabular form (see table I-1) in units of  $10^{-22}$  watts/M $^2$ /c/s. Burst phenomena are measured above this level and are given in terms especially suitable for the variations

observed on this frequency. The basis for the classifications is described by Covington - J.R. Astro. Soc. Can. 45, 49, 1951 and Dodson, Hedeman and Covington, Ap. J. 119, 541, 1954. A modification in terminology with a view to simplification has been introduced and consists essentially of the omission of the descriptive word "Single" from the "Single-Simple" and "Single-Complex" classes; in designating the "Single", "Single-Simple" and "Rise and Fall" bursts into a single classification designated as "Simple Bursts" with an appropriate type number; in the addition of the letter "f" to indicate that the burst deviates from the basic pattern by the presence of one or more small fluctuations in intensity; and by the addition of the letter "A" to indicate that the event has another smaller duration event superimposed upon it.

### Simple Burst

Any single burst which rises to one maximum and then decreases to the pre-burst level.

- $1 \underline{\text{Simple 1}}$  -- Simple burst, type 1 (formerly "single"). Bursts of intensity less than 7 1/2 flux units and duration less than 7 1/2 minutes.
- $2 \frac{\text{Simple 2}}{\text{Bursts of impulsive nature with intensity greater}}$  than  $7 \frac{1}{2}$  flux units.
- $3 \underline{\text{Simple 3}}$  -- Simple burst, type 3 (formerly "rise and fall"). Bursts of moderate intensity with duration greater than 7 1/2 minutes.
- 4 Post-burst increase -- Postburst level is greater than the preburst level. The gradual return to normal flux may require as long as several hours.
  - 5 Absorption following burst (negative post).
- 6 Complex -- (formerly "single-complex"). A single burst which shows two or more comparable maxima before the activity has declined to zero.
- 7 Period of irregular activity or fluctuations -- Series of overlapping bursts of moderate intensity and duration.
- 8 Group -- Series of single isolated bursts occurring in succession with intensity between the events equal to the level before and after the group.
- 9 <u>Precursor</u> -- A small increase of intensity occurring before a larger increase.

### Great Burst

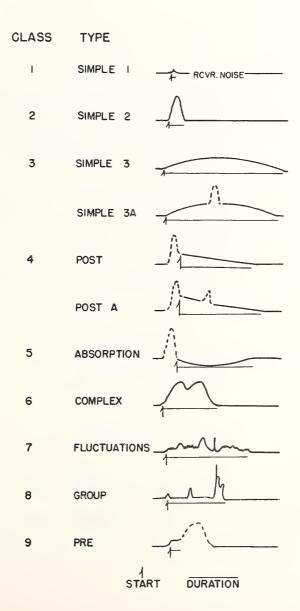
Infrequently occurring bursts of great intensity, often of complicated structure.

### Letter "A"

Indicates that this event has another event superimposed upon it.

### Letter "f"

Indicates that the basic form of the event is modified by secondary fluctuations.



### 200 Mc Observations

Data on solar radio waves made at Cornell University, Ithaca, N.Y. (Marshall Cohen) on 201.5 Mc are presented. All times are in Universal Time (UT or GCT). The antenna is linearly polarized and has a pattern appreciably broader than the solar disk. Flux is reported in units of  $10^{-22}$  watts/m<sup>2</sup>/cps and the tabulated numbers are twice the values observed in the one linear component.

Tables of flux and outstanding occurrences are given in general according to the systems used for the NBS 170 Mc and 450 Mc data.

### 170 Mc and 450 Mc Observations

Data on solar radio emission at the nominal frequencies of 170 Mc and 450 Mc recorded at the Gunbarrel Hill (Boulder) station of the National Bureau of Standards (R.S. Lawrence) are presented. The half width of the antenna lobe is appreciably greater than the solar disk. Polarization is not determined, but the dipole is oriented E-W. All times are in Universal Time (UT or GCT).

3-Hourly and Daily Flux Density and Variability -- Flux density is given in power units. These units are approximately  $10^{-22}$  watts meter- $2(c/s)^{-1}$  for both polarizations together. They will be subject to a correction factor when gain measurements of the antenna have been made. The median flux is measured for every one-hour period having at least thirty minutes of usable record and an applicable gain calibration. A three-hour value of flux is obtained by averaging the available one-hour medians (at least two required). A daily value of flux is obtained by averaging all available one-hour medians (at least four required). A dash indicates that insufficient measurements were made to meet the above requirements or that the records were not of usable quality. Flux values may be followed by the qualifying symbols D, S, and X defined subsequently.

The variability index, given for each three-hour interval, is on a scale 0 to 3 defined as follows:

- 0 The instantaneous flux did not drop below one-half the median level or exceed twice the median level at any time.
  - 1 The instantaneous flux made from one to ten excursions

outside the range described above.

- 2 The instantaneous flux made from ten to one hundred excursions outside the range described above.
- 3 The instantaneous flux made more than one hundred excursions outside the range described above.

For the purpose of the variability index, an excursion whose maximum intensity is M times the median level is counted as M excursions. The variability index is omitted if measurements were made for less than one hour during the period. The variability for the day is the mean of the three-hourly values. The letter S follows variability indices which are in doubt because of atmospherics or local interference.

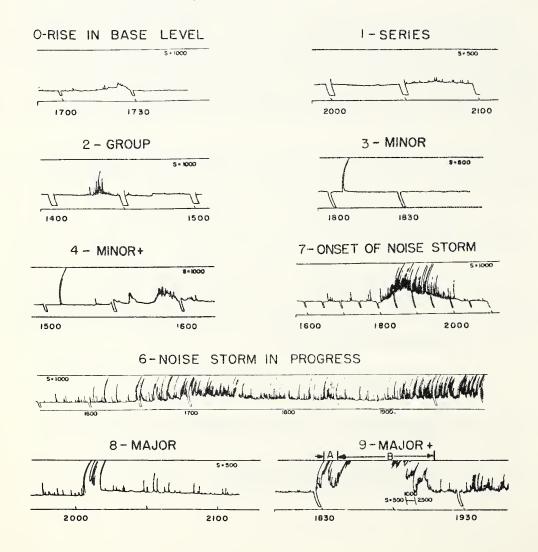
The observing periods are given in U. T. to the nearest 1/10 hour and they usually extend into the next Greenwich day.

Outstanding Occurrences -- A separate table lists the occurrences which are not adequately described by the three-hourly values of flux density and variability. Two classifications are given: (1) A system in general accord with that described and illustrated by Dodson, Hedeman, and Owren (Ap. J. 118, 169, 1953) and (2) the system described in the IGY Solar Activity Instruction Manual, prepared by the Radio Emission editor of the I.A.U. Quarterly Bulletin on Solar Activity.

In system (1) the occurrences are identified by numbers which do not necessarily indicate the magnitude of the event, as follows:

- O Rise in base level -- A temporary increase in the continuum with duration of the order of tens of minutes to an hour.
- l <u>Series of bursts</u> -- Bursts or groups of bursts, occurring intermittently over an interval of time of the order of minutes or hours. Such series of bursts are assigned as distinctive events only when they occur on a smooth record or show as a distinct change in the activity.
- 2 Groups of bursts -- A cluster of bursts occurring in an interval of time of the order of minutes.
- 3 Minor burst -- A burst of moderate or small amplitude, and duration of the order of one or two minutes.
- 4 Minor burst and second part -- A double rise in flux in which the early rise is a minor burst.

- 6 Noise storm -- A temporary increase in radiation characterized by numerous closely spaced bursts, by an increase in the continuum, or by both. Duration is of the order of hours or days.
- 7 Noise storm begins -- The onset of a noise storm occurs at some time during the observing period.
- 8 <u>Major burst</u> -- An outburst, or other burst of large amplitude and more than average duration. A major burst is usually complex, with a duration of the order of one to ten minutes.
- 9A, 9B, or 9 Major burst and second part or large event without distinct first and second parts -- If there is a double rise in flux, the first part, a major burst, is listed as 9A and the second part as 9B. The second part may consist of a rise in base level, a group or series of bursts, a noise storm. A major increase in flux with duration greater than ten minutes but without distinct first and second parts, is listed simply as 9.



In system (2) combinations of the following letters are used to describe some distinctive characteristics of the recorded disturbances:

S = simple rise and fall of intensity,

C = complex variation of intensity,

A = appears to be part of general activity,

D = distinct from (i.e. apparently superimposed upon) the general background,

M = multiple peaks separated by relatively long periods of quietness,

F = multiple peaks separated by relatively short periods of quietness.

E = sudden commencement or rise of activity.

Starting and maximum times are read to the nearest 1/10 minute if they are very definite and otherwise to the nearest minute. If the duration is less than five minutes, it is given to the nearest 1/10 minute; otherwise to the nearest minute (see also qualifying symbols below).

Maximum flux densities are given in units of  $10^{-22}$  watts meter-2(c/s)-1. The instantaneous maximum flux density is the highest peak in the disturbance measured above the sky level. The smoothed maximum flux density is the maximum value of a smooth curve drawn through the outstanding occurrence with a smoothing period of 20 to 50 percent of the total duration; it is measured above the estimated level in the absence of the disturbance. The intention is that (smoothed maximum) x (duration) should give a measure of the energy radiated in the disturbance.

A dash indicates missing or insignificant data. Observations are interrupted during the period from 26 to 29 minutes after each hour for calibrations. Observing periods are given in the Daily Data tables. The following qualifying symbols are used:

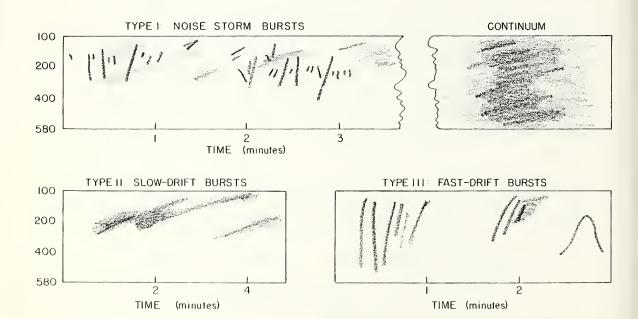
- B Event in progress before observations began.
- D Greater than.
- I Event apparently continued during an interruption of the observations. The period of the interruption may be given in the remarks.
- N See footnotes.
- X Measurement is uncertain or doubtful.
- S Measurement may be influenced by interference or atmospherics.

### Spectrum Observations

Data on solar radio emission in the spectral range 100-580 Mc recorded at the Harvard University Radio Astronomy Station, Fort Davis, Texas (A. Maxwell) are presented. The research is sponsored by the Geophysics Research Directorate of the Air Force Cambridge Research Center, Air Research and Development Command, under contract AF19(604)-1394.

The receiving equipment consists of three separate sweep-frequency receivers covering the bands 100-180, 160-320, 300-580 Mc. These are attached to separate broad-band feeds mounted coaxially at the primary focus of an 8.55 meter diameter paraboloid, the 160-320 Mc feed being cross-polarized with the other two feeds. The effective collecting area of the antenna is 40 sq. meters at 100 Mc and 45 sq. meters at 500 Mc.

The four types of recognized spectral activity are idealized below:



Type IV continuum radiation is a steady enhancement of the back-ground level over a wide band of the spectrum. In one form it is frequently associated with noise storms. A second form is characterized by the following properties:

(1) It is uniformly distributed over a band of frequencies often as wide as 300 Mc. The whole band may drift systematically toward higher or lower frequencies.

(2) Its intensity is essentially non-fluctuating.

(3) It is usually of high intensity, i.e., greater than  $10^{-20}$  watts  $meter^{-2}(c/s)^{-1}$ .

(4) It often occurs at frequencies higher than the spectral range of noise storms, the upper limit of which rarely exceeds 250 Mc.

(5) After great radio outbursts it may last for as long as 5 hours. At the other extreme, a miniscule version, occurring after a group of fast drift bursts or an inverted U burst, may last only 10-60 seconds.

The large scale examples of this continuum are listed as "Cont. IV" in the tables. It probably corresponds to the "Type IV" radiation described by Boischot (Comptes Rendus 244, 1326, 1957) from fixed frequency observations taken at 169 Mc at Meudon, France. Photographic examples are published by Maxwell, Swarup and Thompson (Proc. IRE 46, 142, 1958). A few remaining solar radio bursts are tabulated as unclassified.

The symbols used in the tables are:

b = single burst

g = small group (<10) of bursts

G = large group (>10) of bursts

= Arrows indicate continuity of solar activity between two Greenwich days.

The minimum detectable level of solar activity is a function of frequency: approximately  $5 \times 10^{-22}$  watts meter- $2(c/s)^{-1}$  at 100 Mc and  $10^{-21}$  watts meter- $2(c/s)^{-1}$  at 500 Mc. The equipment records signals over an intensity range of approximately 1000:1. There are three classes of intensity given in the tables. For 100 Mc they are:

1 = faint, 5 to 30 x  $10^{-22}$  watts meter  $-2(c/s)^{-1}$ 

2 = moderate, 30 to 100 x  $10^{-22}$ 

3 = strong, >100 x 10-22.

The times are Universal Time (UT). The accuracy is to the nearest half minute, except in the case of major outbursts which are specified to the nearest 0.1 minute.

C. Kp. Ap. and Selected Quiet and Disturbed Days -- The data in the table are: (1) preliminary international character figures, C; (2) geomagnetic planetary three-hour range indices, Kp; (3) daily "equivalent amplitude," Ap; (4) magnetically selected quiet and disturbed days.

This table is made available by the Committee on Characterization of Magnetic Disturbance of IAGA, IUGG. The Meteorological Office, De Bilt, Holland collects the data from magnetic observatories distributed throughout the world, and compiles C and selected days. The Chairman of the Committee computes the planetary and equivalent amplitude indices. The same data are also published quarterly in the Journal of Geophysical Research along with data on sudden commencements (sc) and solar flare effects (sfe).

The C-figure is the arithmetic mean of the subjective classification by all observatories of each day's magnetic activity on a scale of O (quiet) to 2 (storm).

Kp is the mean standardized K-index from 12 observatories between geomagnetic latitudes 47 and 63 degrees. The scale is 0 (very quiet) to 9 (extremely disturbed), expressed in thirds of a unit, e.g. 5- is 4 2/3, 50 is 5 0/3, and 5+ is 5 1/3. This planetary index is designed to measure solar particle-radiation by its magnetic effects, specifically to meet the needs of research workers in the ionospheric field. A complete description of Kp has appeared in Bulletin 12b, "Geomagnetic Indices C and K, 1948" of the Association of Terrestrial Magnetism and Electricity (IATME), International Union of Geodesy and Geophysics.

Ap is a daily index of magnetic activity on a linear scale rather than on the quasi-logarithmic scale of the K-indices. It is the average of the eight values of an intermediate 3-hourly index "ap," defined as one-half the average gamma range of the most disturbed of the three force components, in the three-hour interval at standard stations; in practice, ap is computed from the Kp for the 3-hour interval. The extreme range of the scale of Ap is 0 to 400. The method is described in IATME Bulletin No. 12h (for 1953) p. viii f. Values of Ap (like Kp and Cp) have been published for the Polar Year 1932/33 and for the years 1937 onwards.

The magnetically quiet and disturbed days are selected in accordance with the general outline in Terr. Maq. (predecessor to J. Geophys. Res.) 48, pp 219-227, December 1943. The method in current use calls for ranking the days of a month by their geomagnetic activity as determined from the following three criteria with equal weight: (1) the sum of the eight Kp's; (2) the sum of the squares of the eight Kp's; and (3) the greatest Kp.

<u>Chart of Kp by Solar Rotations</u> -- The graph of Kp by solar rotations is furnished through the courtesy of Dr. J. Bartels, Geophysikalisches Institute, Göttingen.

### VI RADIO PROPAGATION QUALITY INDICES

One can take as the definition of a radio propagation quality index: the measure of the efficiency of a medium-powered radio circuit operated under ideal conditions in all respects, except for the variable effect of the ionosphere on the propagation of the transmitted signal. The indices given here are derived from monitoring and circuit performance reports, and are the nearest practical approximation to the ideal index of propagation quality.

Quality indices are usually expressed on a scale that ranges from one to nine. Indices of four or less are generally taken to represent significant disturbance. (Note that for geomagnetic K-indices, disturbance is represented by higher numbers.) The adjectival equivalents of the integral quality indices are as follows:

1 = useless 4 = poor-to-fair 7 = good 2 = very poor 5 = fair 8 = very good3 = poor 6 = fair-to-good 9 = excellent

CRPL forecasts are expressed on the same scale. The tables summarizing the outcome of forecasts include categories P-Perfect; S-Satisfactory; U-Unsatisfactory; F-Failure. The following conventions apply:

- P forecast quality equal to observed U forecast quality two or more grades different from observed when both forecast and observed were > 5, or both < 5
- S forecast quality one grade F other times when forecast quality two or more grades different from observed

Full discussion of the reliability of forecasts requires consideration of many factors besides the over-simplified summary given.

The quality figures represent a consensus of experience with radio propagation conditions. Since they are based entirely on monitoring or traffic reports, the reasons for low quality are not necessarily known and may not be limited to ionospheric storminess. For instance, low quality may result from improper frequency usage for the path and time of day. Although, wherever it is reported, frequency usage is included in the rating of reports, it must often

be an assumption that the reports refer to optimum working frequencies. It is more difficult to eliminate from the indices conditions of low quality for reasons such as multipath or interference. These considerations should be taken into account in interpreting research correlations between the Q-figures and solar, auroral, geomagnetic or similar indices.

North Atlantic Radio Path -- The CRPL quality figures, Qa, are compiled by the North Atlantic Radio Warning Service (NARWS), the CRPL forecasting center at Ft. Belvoir, Virginia, from radio traffic data for North Atlantic transmission paths closely approximating New York-to-London. These are reported to CRPL by the Canadian Defense Research Board, Canadian Broadcasting Corporation, and the following agencies of the U. S. Government:--Coast Guard, Navy, Army Signal Corps, U. S. Information Agency. Supplementing these data are CRPL monitoring, direction-finding observations and field-strength measurements of North Atlantic transmissions made at Belvoir.

The original reports are submitted on various scales and for various time intervals. The observations for each 6-hour interval are averaged on the original scale. These 6-hour indices are then adjusted to the 1 to 9 quality-figure scale by a conversion table prepared by comparing the distribution of these indices for at least four months, usually a year, with a master distribution determined from analysis of the reports originally made on the 1 to 9 quality-figure scale. A report whose distribution is the same as the master is thereby converted linearly to the Q-figure scale. The 6-hourly quality figure is the mean of the reports available for that period.

The 6-hourly quality figures are given in this table to the nearest one-third of a unit, e.g. 50 is 5 and 0/3; 5- is 4 and 2/3; 5+ is 5 and 1/3. Other data included are:

- (a) Whole-day radio quality indices, which are weighted averages of the four 6-hourly indices, with half weight given to quality grades 5 and 6. This procedure tends to give whole-day indices suitable for comparison with whole-day advance forecasts which seek to designate the days of significant disturbance or unusually quiet conditions.
- (b) Short-term forecasts, issued every six hours by the North Atlantic Radio Warning Service. These are issued one hour before  $00^h$ ,  $06^h$ ,  $12^h$ ,  $18^h$ , UT and are applicable to the period 1 to 7 hours ahead.
- (c) Advance forecasts, issued twice weekly by the NARWS (CRPL-J reports) and applicable 1 to 3 or 4 days ahead, 4 or 5 to 7 days ahead, and 8 to 25 days ahead. These forecasts are scored against the whole-day quality indices.

(d) Half-day averages of the geomagnetic K indices measured by the Fredericksburg Magnetic Observatory of the U. S. Coast and Geodetic Survey.

A chart compares the short-term forecasts with Qa-figures. A second chart compares the outcome of advance forecasts (1 to 3 or 4 days ahead) with a type of "blind" forecast. For the latter, the frequency for each quality grade, as determined from the distribution of quality grades in the four most recent months of the current season, is partitioned among the grades observed in the current month in proportion to the frequencies observed in the current month.

Ranges of useful frequencies on the North Atlantic radio path are shown in a series of diagrams, one for each day. The shaded area indicates the range of frequencies for which transmissions of quality 5 or greater were observed. The blacker the diagram, the quieter the day has been; a narrow strip indicates either high LUHF, low MUF, or both. These diagrams are based on data reported to CRPL by the German Post Office through the Fermeldetechnischen Zentralamtes, Darmstadt, Germany, being observations every one and a half hours of selected transmitters located in the eastern portion of North America. Since January 6, 1958 the transmitters monitored are restricted to those located north of 39° latitude. The magnetic activity index, AFr, from Fredericksburg, Va., is also given for each day.

Note: Beginning with data for September 1955, Qa has been determined from reports that are available within a few hours or at most within a few days, including for the first time, the CRPL observations. Therefore these are the indices by which the forecasters assess every day the conditions in the recent past. Over a period of several years, they have closely paralleled the former Qa indices which excluded CRPL observations and included three additional reports received after a considerable lag. Qa was first published to the nearest one-third of a unit at the same time.

North Pacific Radio Path -- The CRPL quality figures, Qp, are compiled by the North Pacific Radio Warning Service (NPRWS), the CRPL forecasting center at Anchorage, Alaska, from radio traffic data for moderately long transmission paths in the North Pacific equivalent to Seattle-to-Anchorage or Anchorage-to-Tokyo. These include reports to CRPL by the Alaska Communications System, Aeronautical Radio, Inc., U. S. Air Force and Civil Aeronautical Administration. In addition, there are CRPL monitoring, direction finder observations and field strength measurements of suitable transmissions.

The original reports are on various scales and for various time intervals. The observations for each 8 hours or 24 hour period are averaged on the original scale. This average is compared with reports for the same period in the preceding two months and expressed

as a deviation from the 3-month mean. The deviations are put on the 1 to 9 scale of quality which is assumed to have a standard deviation of 1.25 and a mean for the various periods as follows:

03-10	hours	UT	5.33
11-18			5.33
19-02			6.00
00-24			5.67

The 8-hour and 24-hour indices Qp are determined separately. Each index is a weighted mean where the CRPL observations have unit weight and the others are weighted by the correlation coefficient with the CRPL observations.

The table, analogous to that for Qa, includes the 8-hourly quality figures; whole day quality figures; short-term forecasts issued by NPRWS three times daily at  $02^h$ ,  $10^h$ , and  $18^h$  UT, applicable to the stated 8-hour periods; advance forecasts issued twice weekly by NPRWS (CRPL-Jp report); and half-day averages of geomagnetic K indices from Sitka.

The chart compares the outcome of advance forecasts, on the same basis as the similar chart for the North Atlantic Radio Path.

Note: Beginning with November 1956 the short-term forecast formerly made at 0900 UT was changed to 1000 UT. The North Pacific quality figures used for evaluation are now 8-hourly rather than 9-hourly.

### VII ALERT PERIODS AND SPECIAL WORLD INTERVALS

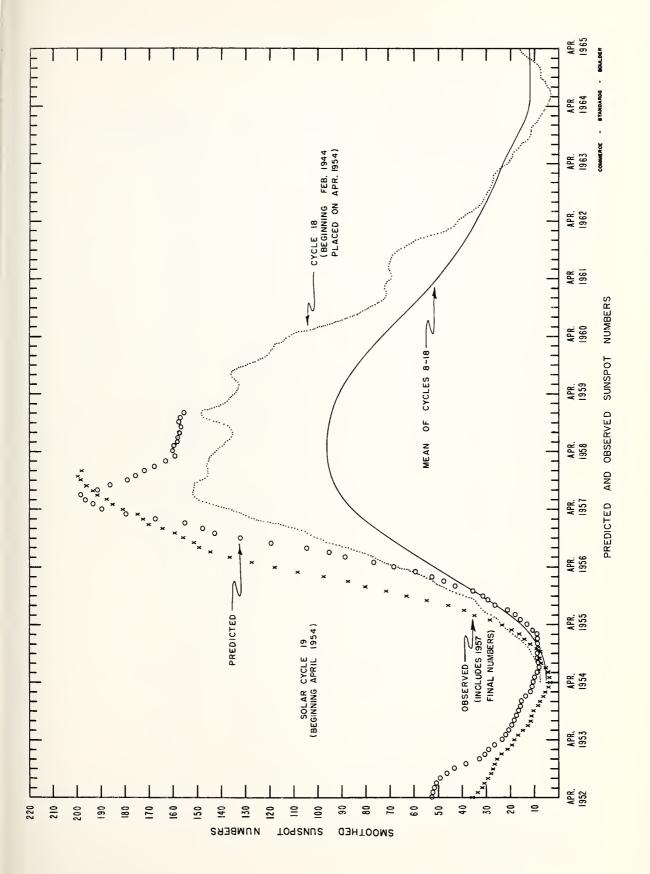
A table gives the Alert Periods and Special World Intervals (SWI) as designated by the IGY World Warning Agency at Ft. Belvoir, Va. For each day of the Alert or SWI are given the number of flares of importance two or greater reported promptly to the IGY World Warning Agency and the magnetic activity index  ${\rm A}_{Be}$  observed at the IGY World Warning Agency.



May 1958	American Relative Sunspot Numbers R <sub>A</sub> ,
1	199
2	211
3	265
4	240
5	212
6	172
7	183
8	153
9	176
10	180
11	188
12	148
13	152
14	107
15	120
16	117
17	120
18	130
19	159
20	161
21	171
22	158
23	197
24	206
25	175
26	173
27	145
28	153
29	189
30	179
31	190
Mean:	171.9

June 1958	Zürich Provisional Relative Sunspot Numbers R <sub>Z</sub>	Daily Values Solar Flux at 2800 Mc, Ottawa, Canada Flux
1	200	<b>21</b> 9
2	154	220
3	181	227
4	195	246
5	195	256
6	176	260
7	185	238
8	200	233
9	209	252
10	200	234
11	193	235
12	193	227
13	176	220
14	160	208
15	131	197
16	100	191
17	113	182
18	100	177
19	114	189
20	107	193
21	141	194
22	148	213
23	184	217
24	189	221
25	199	226
26 27 28 29 30 31	183 178 174 200 159	233 237 232 220 217
Mean:	167.9	220.5

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### CALCIUM PLAGE AND SUNSPOT REGIONS JUNE 1958

CMP		McMath	Return	Ca	lcium P	lage Data	Sunsp	oot Data
June	Lat	Plage	of	CMP V			CMP Value	
1958		Number	Region	Area	Int.	History, Age	Area Cour	
01.5 02.6 02.7 03.6 04.4	N31 N08 N22 S13 N26	4588 4582 4583 4585 4587	+ 4539 ++ New 4538	2200 800 3800 (1300) 2200	2 1.5 3 (2.5) 2	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	(20) (2) 320 8 440 1 110 2	b — d ℓ√ℓ ℓ — ℓ b — d
05.1	S22	4589	New	800	2	$\ell \smile \ell - 1$	100 2	l ~ d
06.3	N18	4591	4542,43	5200	3	l - l 2,4	100 6	ℓ - d
09.3	S18	4592	New	4600	3.5	$\ell \sim \ell$ 1	400 1	$\ell \frown \ell$
10.0	N28	4596	New	10,000	3.5	l → / l 1	1870 16	l ~ l
10.2 10.4 11.6 11.9 12.7	N41 N16 S20 S11 N18	4597 4599 4598 4600 4601	New New 4548 4548 4552	4000 1400 5000 2500 3000	3.5 3 3 3	\( \lambda - \ell \) 1 \( \ell \ - \ell \) 1 \( \ell - \ell \) 2 \( \ell - \ell \) 2 \( \ell \ \times \ell \) 4	930 8 20 2 70 3 10 7 60 2	ℓ ∧ ℓ b − d ℓ \ d b − d b − d
14.3 14.7 15.2 15.5 16.6	S14 N24 S11 N26 S24	4602 4603 4605 4611 4604	+++ 4556 4555 ** +++	800 1000 500 500 2700	1.5 2 1.5 1	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	(10) (1) (10) (1) 10 1	
16.7 17.9 18.9 19.6 19.8	S13 N12 S24 N40 N13	4606 4607 4608 4609 4610	4555 4563 New New 4561	1200 5400 4000 1800 1600	2 3 3 2 2.5	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	820 6 290 2 80 5	l ~ l l ~ l b ~ l
21.6 22.1 22.7 23.0 24.6	N26 N07 N19 S18 N14	4612 4614 4613 4615 4616	4560 4575 4568 4576 4574,77	1600 600 1300 1000 5000	2.5 2 2 1 3.5	$ \begin{array}{ccccc} l & - & l & 3 \\ l & \nabla & l & 2 \\ l & \sim & l & 6 \\ l & \nabla & l & 4 \\ l & - & l & 3 \end{array} $	70 3 110 7 780 11	l \ l l \ d
26.0 26.4 27.7 27.7 27.8	N25 S16 N28 N18 N08	4617 4618 4619 4621 4620	4574,77 * 4578,88 4578 New	2100 11,000 4500 2600 1000	2.5 3 2.5 3 2.5	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	190 10 (50) (3)	l\ a
29.0 29.7 29.9	S20 N11 S08	4622 4623 4624	New 4582,83 New	6000 4500 2700	3.5 3 3.5	$ \begin{array}{cccc} \ell & -\ell & 1 \\ \ell & \sim \ell & 3 \\ \ell & -\ell & 1 \end{array} $	1180 14 300 7 220 9	$ \begin{array}{c} \ell \sim \ell \\ \ell \sim \ell \\ \ell - d \end{array} $

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<sup>+</sup> = Remnants of 4531.

<sup>++</sup> = 4541 and part of 4537.

<sup>+++</sup> = Remnant of 4553.

<sup>++++</sup> = 4565 and part of 4553.

<sup>\* = 4579, 4580, 4584.</sup> 

<sup>\*\* =</sup> In position of 4556.

### CORONAL LINE EMISSION INDICES JUNE 1958

																								_			_				
int iter)	R <sub>1</sub>	×	×	×	×	×	×	×	78a	150a	×	×	×	×	×	54а	×	×	27	09	×	×	×	×	×	×	×	×	×	81	150
North West Quadrant (observed 7 days later)	R <sub>6</sub>	×	×	×	×	×	×	×	47a	75a	×	×	×	×	×	28a	×	×	53	07	×	×	×	×	×	×	×	×	×	7.7	58
rth West erved 7	$_{1}^{G_{1}}$	×	191	×	×	221a	230a	272a	209в	250a	150a	×	×	×	×	113a	100g	150a	186	194	180a	×	×	×	×	×	×	200a	×	210	260
No)	99	×	140	×	×	161a	160a	180a	160a	220a	131a	×	×	×	×	83a	85a	150a	125	176	129a	×	×	×	×	×	×	156a	×	153	526
ant iter)	$^{R_1}$	×	×	×	×	×	×	×	96 <b>a</b>	×	×	×	×	×	×	50a	×	×	×	57	×	×	×	×	×	×	×	×	×	7,7	53
South West Quadrant (observed 7 days later)	$^{R}6$	×	×	×	×	×	×	×	38a	×	×	×	×	×	×	22a	×	×	×	18	×	×	×	×	×	×	×	×	×	3%	77
uth Wes rved 7	$^{\mathrm{l}_{\mathrm{9}}}$	×	89	×	×	112a	96a	128a	220a	×	150a	. ×	×	×	×	76a	135a	150a	107	×	×	×	×	×	×	×	×	140a	×	72	136
So Sobse	$e^{6}$	×	57	×	×	68a	82a	988	126a	×	112a	×	×	×	×	51a	92a	93a	72	×	×	×	×	×	×	×	×	95a	×	74	77
nt Lier)	$^{R_1}$	×	76a	×	×	×	×	×	×	×	×	×	×	×	×	×	×	×	×	×	×	×	50a	, S	×	×	×	×	×	55a	×
South East Quadrant (observed 7 days earlier)	R6	×	31a	×	×	×	×	×	×	×	×	×	×	×	×	×	×	×	×	×	×	×	20g	27a	×	×	×	×	×	37a	×
th East red 7 da	$G_1$	68	×	×	×	в†/6	×	×	×	×	151a	×	×	×	×	×	143	300g	×	80	71a	06	140a		125a	×	×	×	×	260a	145a
Sou (obser	99	99	×	×	×	77a	×	×	×	×	117a	×	×	×	×	×	8	204a	×	63	62a	22	93a	70a	85a	×	×	×	×	202a	92a
rant arlier)	$^{\mathrm{R}_{1}}$	×	т94	×	×	×	×	×	×	×	×	×	×	×	×	×	×	×	×	×	×	× 6	g (	40a	×	×	×	×	×	72a	×
	.R	×	31a	×	×	×	×	×	×	×	×	×	×	×	×	×	×	×	×	×	×	× :	44a	2/8	×	×	×	×	×	45а	×
North East Quad (observed 7 days e	$_{1}^{G_{1}}$	178	×	336a	×	320a	×	×	×	×	116a	×	×	×	×	×	25	250a	× 8	3:	144a	252	220a	130a	210a	×	×	×	×	306a	180a
Nort (observ	e <sub>o</sub>				_	233a		_	_			_		_	_		_				_	_	_	-	_				_	-	
CMP	1958	П	2	~	7	٠.	9		∞	6	10	11	27	13	77	15	16	17	18	19	50	21	22	23	24	25	56	27	28	56	30

# = yellow line observed.
a = index computed from low weight data.
x = no observations.

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### SOLAR FLARES JUNE 1958

PROVISIONAL IONOSPHERIC Slow S-SWP EFFECT S-SWP G-SWP G-SWF S-SWP S-SWF PAGE 80 72 18 125 28 92 50 58 106 28 146 58 64 MAX. 2.00 2.10 2.10 1.00 3.54 2 • 30 2.45 69.0 1.90 MAX. WIDTH Ha 2.40 3.10 3.00 5.00 6.00 2.31 8.44 6.10 14.00 2.60 1.00 4.80 8.60 10.00 2.30 2.61 4.97 5.12 CORR. AREA Sq. Deg 2.20 3.65 2.90 1.60 .62 2.00 2.50 1.50 1.20 2.00 2.20 5.40 4.22 3.30 7.00 3.60 4.50 2.00 1.84 .91 MEAS. AREA Sq. Deg. 1042 1113 1120 1120 1233 1243 1314 1745 1913 2119 0404 091**6** 1210 1512 1514 1515 1516 1740 2126 0937 1155 1228 1022 1410 1504 1504 1537 0000 0859 0905 1353 TDATE U T OBS. 2221 12 N N ~ 464 12000 N P N N V V V 2 m N m 22 IM. POR-00000 000 00 0 0000 200 00000 0 00000000 DURA-TION -MINUTES 114 123 123 146 150 160 170 170 27 25 25 25 25 25 25 25 25 30 30 20 22 22 24 26 11 10 15632296 4537 4587 4587 4587 4578 4574 4578 4578 4578 4578, 4578 4578 4578 4578 4578 4578 4578 4578 4578 45978 45978 McMATH PLAGE REGION LOCATION #10 #46 #46 #51 E26 E27 2337 2337 2350 2450 2457 2457 MER. DIST. APPROX. LAT. 2119 2120 2126 MAX. 1014 1514 0937 406U 1501 2152 UNIVERSAL TIME 2 2  $\bigcirc$ 00 222 OBSERVED 1110 1146 1140 1126 1402 1400 1309 1321 1321 1748 1952 2135 2135 0527 0710 0917 0938 0913 0925 0946 1235 1235 1345 1428 1518 1518 1513 1645 0000 END تا يا يا تا ш Ш च व्य out to build build 9.1 1037 1103 1103 11107 11112 11333 1339 1242 1242 1315 1743 1907 2106 2107 2107 START 0335. 0400 0914 1202. 1320 1507 1511 1512 1512 0 9 3 2 1 1 4 3 1 1 2 2 8 1 3 2 2 1 4 5 1 1 4 5 7 1 4 5 7 1 6 2 9 2 1 4 7 00000 0258 0520 0700 0841 0849 0850 0850 1958 JUNE \*\*\*\*\*\*\*\*\*\*\*\*\* SYDNEY
CHITAKA
ONDREJOV
HCKARL
NEDERHORST
GAC PEAK
HCMAH
R O HERST
CAPRI S
COPRI S
CORRI S GOOD HOPE GOOD HOPE CAPRI S ONDREJOV ZURICH WENDEL GOOD HOPE
ARCETRI
ARCETRI
20RICH
ZURICH
GSNRL
CAPRI S
STUCKHOLM
STUCKHOLM
SAC PEAK AITAKA
TASHKENT
NIZARNAH
LOCARNO
CAPRI S
SCHAUINS
NIZARIAH
NEDERHORST
SSTOCKHOLM
Z SUTCKHOLM
R O HERST WENDEL ZURICH ONDREJOV ONDREJOV USNRL ACMATH SAC PEAK OBSERVATORY

COMMERCE - STANDANDS - BICKLINGS

17	Di Bic				_		ΜE	M.F						WF																												
PROVISIONAL	IONOSPHERIC	Preci				3MC-C	Slow S-SWF	Slow S-SWP			G-SWF	S-SWF		Slow S-SWP									# n			S-SWF																PAGE 2
	MAX.	*		124	28	120	£47	126	149	105	100	107													-	99	,	2	7 2	7.2											100	6 58
	MAX.	Жа		2,00		2.00	00.7	7.00	1.80	1.85	1.68	50.5		3.00	7.00				_												5.20	2.10	09.7	,			2 • 30			2.50		
MEASUREMENTS	CORR.	Bq. Deg.	3.00	3.60	1	5.20	0 • •	3.26	24.80	1.94	1.23	1.23		4 . 83	•	11.00	8	4.00	00.0		00.6	7.44	9 00	2.00	2.00	12.00	3.00	3.00	3.17	2.54				3.20	4 • 80	2 - 80		60.6	2.50		1.13	000
	MEAS.	Sq. Deg.		3.60	6.70	2.03	79.7	1.02	11.40	68.	60.	68.		1.21	0 0							1.28				1.63	)	1.70	77	1.46				2.20	2.50	2.00	!		000.1		.82	00.
	TIME	T D	0944	1502	1	1628	11/1	1818	2310	0017	0112	55.10		0448	0000							1540				1622		1708	1245	2:18	0642	0655	0470	0 725	6260	1042	1026		1,12	1109	1309	1761
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	LAT		327	510	\$18	518	221	2 1 2	146	N46	716	2 Z Z	1001	N16	72	142	N 26	243	N17	240	N43	N45	2 5 5	1 18	1427	2 4 Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z	113	541	516	842	43	127	43	129	516	128	:26	.127	127	₩26		T 7
		PHASE		1530	1635	1628	77/7	1818	2310				0439	0448								1338				1622		1708	1836	2118	0.042	0655				1.44			7177	1109	1509	1771
OBSERVED	UNIVERSAL TIME END		0360	1524	1815	1658	1001 1745 D		2330	0025	0124	1410	0518	0503	0507 D		0648 U		0920 D		1110 D					1633	1642	1719	1845	2129 D	0645	0657	0136	0928 D	0928 D	1025	1044	1133	1125 D		1316	10
	START		9 0460	1458 E	1615	1619	1728 6		2300	0017	0112 E		0436		0458 E		0620	0825		1021 E	1056		1529 F				1618 =	1705	1835	2100	0638		0819			1019	1026 E		1102 1105 E		1300	2101
DATE	1958	DUNE	0.5	0 C	0 0 2	0 0	2 0	000	60	90	90	900	90	90	90	90	0 0	90	90	900	90	90	90	90	90	90	90	90	90	90	0.7	07	0 2	0.7	07	2 0	0.7	07	2 0	0.7	27	5
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28	
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12	

PROVISIONAL	EFFECT							Slow S-SWF			S. CILE	JMC-C													AMS-S		S-SWF														S-SWF					C U
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MEAS.	AREA Sq. Deg.		• 8 J				2.40	2.50			1.50	1.00				1.20	_				-				2.80	,	2.40	3.20	2.60	5.60	07.							0	2.50	)		_	2.30	9		
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### SOLAR FLARES JUNE 1958

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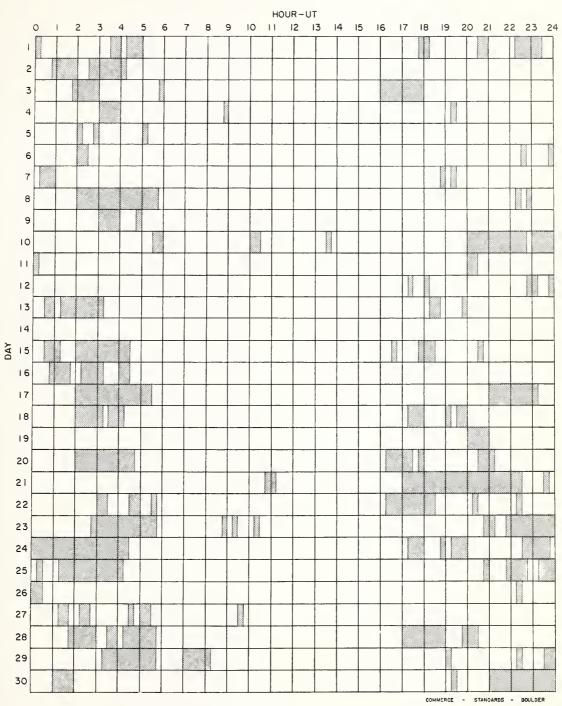
### SOLAR FLARES UNE 1958

JUNE 1958

ANACAPRI SWEDISH
KODAIKANAL
KRASKA PAGHRA
RYAL OBSERVATORY, EDINBURGH
CREENWICH ROYAL OBSERVATORY, HERSTMONCEUX
SACRAMENTO PEAK
SCHAUINSLAND
UNITED STATES NAVAL RESEARCH LABORATORY
ROYAL OBSERVATORY, CAPE OF GOD HOPE CAPRI S
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SAC PEAK: ALL-VALUES IN MAX. INT. COLUMN ARE ARBITRARY UNITS (0-40), NOT PERCENT OF CONTINUOUS SPECTRUM. E - LESS THAN
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### INTERVALS OF NO FLARE PATROL OBSERVATIONS $$_{\rm JUNE~1958}$$



Times indicated are accurate to the nearest 15 minutes.

#### Stations included:

Herstmonceux

Anacapri (Swedish)
Arcetri
Arosa
Athens
Climax
Dunsink
Greenwich Royal Observatory,

Hawaii Huancayo Kodaikanal Locarno Meudon Mitaka Nizamiah Ondrejov Ottawa Royal Observatory, Edinburgh U.S. Naval Research Laboratory Zurich Hik

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#### SUBFLARES NOTEO AS FOLLOWS. DATE - UNIVERSAL TIME - COORDINATES

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# SOLAR FLARES Angust 1957

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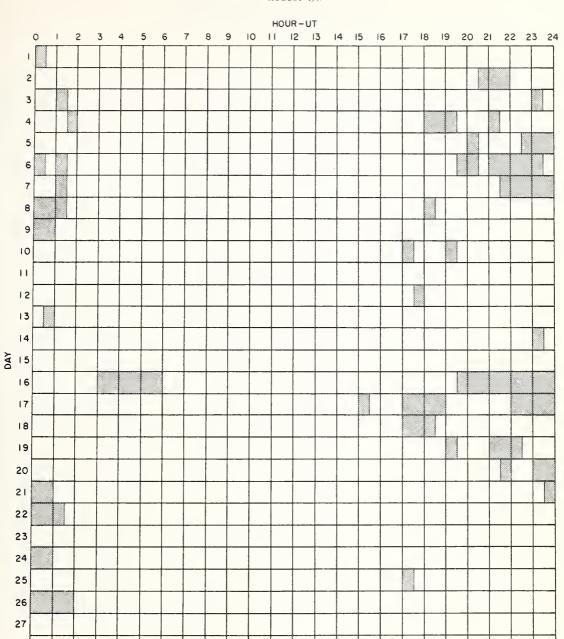
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ORSFRVATORY			UNIVERSAL TIME		APPROX.	1	McMATH	NO.	POR.	COND	TIME	MEAS.	CORR.	MAX.	MAX.	IONOSPHERIC
	1957 AUG	START	END	MAX. PHASE	LAT.	MER. DIST.	PLAGE	MUNUTES	TANCE		L D	AREA Sq. Deg.	AREA Sq. Deg.	WIDTH	INT. %	EFFECT
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These flare reports are addenda to the August 1957 flares published in CRPL-F 157 part B, September 1957.

### INTERVALS OF NO FLARE PATROL OBSERVATIONS

AUGUST 1957



COMMERCE - STANDARDS - BOULDER

Times indicated are accurate to the nearest half hour.

#### Stations included:

Anacapri (Swedish)
Arcetri
Athens
Climax
Dunsink
Greenwich Royal Observatory,
Herstmonceux
Hawaii

Huancayo Kodaikanal Krasnaya Pakhra Meudon Mitaka Nederhorst Nizamiah Ondrejov Ottawa Royal Observatory, Edinburgh Sacramento Peak Simeis Sydney Uccle U.S. Naval Research Laboratory

Utrecht Zurich

## IONOSPHERIC EFFECTS OF SOLAR FLARES (SHORT-WAVE RADIO FADEOUTS) MAY 1958

	Start	End	Туре	Wide	Impor-	Observation Stations	Known
May	UT	UT		Spread	tance		Flare, UT
1958				Index			CRPL-F 166 B
1	0826	0901	S-SWF	1	2	PU	0821
1	1006	1021	S-SWF	3	1+	KU, NE	0954
1	1410	1455	Slow S-SWF	5	2-	BE, CR, HU, MC, NE PR, WS	1353
1	1815	1850	S-SWF	5	3-	AD, BE, CR, HU, MC, PR	1806
1	2130	2155	Slow S-SWF	5	1+	AD, BE, MC, PR, WS	2115
1	2331	2341	S-SWF	3	1+	AN, TO	2327
2	0543	0621	S-SWF	4	1	ко, ок	0546E
2	0755	0818	S-SWF	1	1	NE NE	0754
2	0830	0850	S-SWF	1	1	NE	0819
3	1035	1140	S-SWF	1	3	<u>10</u>	1015E
3	1306	1400	Slow S-SWF	5	2+	BE, HU, MC, NE, PR	1300
3	1611	1635	Slow S-SWF	4	2	JU, MC, WS	1617E
4	0328	0430	S-SWF	5	2	AD, OK, TO, CW+	101,2
4	0743	0801	S-SWF	3	2-	NE, PU	0743E
4	1643	1707	S-SWF	5	2+	BE, CR, HU, MC, NE, PR, WS	1643
5	0011	0030	S-SWF	3	2	AD, OK	0012
5	0203	0234	Slow S-SWF	1	3	OK, OK	0221E
5	0407	0500	S-SWF	5	3	$\frac{\partial K}{\partial D}$ , $OK$ , TO, RCA+, CW $\frac{1}{1}$ +	0407
5	0911	1002	S-SWF	4	2	NE, SW, CW***	0856
5	1019	1045	Slow S-SWF	1	2	NE NE	0945
5	1218	1242	G-SWF	5	1	BE, MC, NE, PR, CW***	1205
5	1326	1405	G-SWF	2	1	MC, PR	1315
5	1928	2012	S-SWF	5	2	BE, CR, HU, MC, PR	1925
5	2032	2110	S-SWF	5	2+	AD, BE, CR, HU, MC, PR, TO, RCA+	2032
6	0210	0233	G-SWF	4	1	AD, OK, TO	0210E
	2212	0110	g: g:=				
6	0340	0419	S-SWF	4	2-	AD, <u>OK</u> , TO, CW+	0325
8	0645	0657	S-SWF	1	2	KO WO	0650E
8	0740 0912	0800 0923	Slow S-SWF	1	1	<u>xo</u> Ju	0837
10	0844	0923	S-SWF S-SWF	1 5	3 2	DA, JU, NE, PU	0910 0844E
10	0044	0920	3-3WF	ر	-	DA, 30, NE, <u>PU</u>	U044E
17	1355	1422	Slow S-SWF	5	2-	BE, HU, MC, NE, PR, PU	1340
17	2135	2205	G-SWF	4	1	AD, BE, MC, WS	2136
19	0425	0507	Slow S-SWF	3	2-	<u>OK</u> , TO	0425E
27	1653	1740	Slow S-SWF	4	1+	HU, MC, PR, WS	
29	0035	0110	Slow S-SWF	3	1	AD, OK	
31	0556	0615	S-SWF	5	1+	NE, OK, PU	0557E

## SOLAR RADIO EMISSION OUTSTANDING OCCURRENCES JUNE 1958

OTTAWA 2800 MC

June 1958	Type*	Start UT Hrs:Mins	Duration Hrs:Mins	Maxim Time UT Hrs:Mins	um Peak Flux	Remarks
2 2 3 3	6 Complex 1 Simple 1 3 Simple 3 A 2 Simple 2 2 Simple 2	19 48.3 22 51.5 15 09 15 09.3 19 33.9	5 2,5 35 6 1	19 50 22 52.5 15 22 15 10.5 19 34.1	8 6 9 215 11	
3 4 4	1 Simple 1 2 Simple 2 Rise**	21 32.7 11 04.5 14 52	1 8	21 33.2 11 05.1	4 45	
4	3 Simple 3 A 6 Complex f	21 25 21 38	1 55 47	indet. 21 56	20 570	
5	3 Simple 3 A 8 Group (2) 2 Simple 2 f	16 14 16 14 16 14	3 1 28 18	indet.	25 380	
	4 Post increase f 6 Complex f 2 Simple 2	16 55 18 09	23 47 4	17 10 18 10	25 330 12	
7 8 8 8 10	6 Complex 6 Complex 1 Simple 1 1 Simple 1 1 Simple 1	10 07.5 16 49.2 16 57.5 21 04.5 12 56.5	1.0 2 2 1 1.5	10 12.5 16 49.9 16 58.2 21 04.8 12 57.2	70 7 3 4 6	
10 11 11	1 Simple 1 2 Simple 2 2 Simple 2 f 4 Post Increase	18 36 11 37.5 13 06	3 3 3•5 2 15	18 37.5 11 38 13 07.5	3 30 145 20	
11 11 11	8 Group (2) 1 Simple 1 2 Simple 2 f 1 Simple 1 6 Complex 4 Post Increase	16 35.2 16 35.2 16 37.6 17 48.7 20 38	4.4 1.2 2 1.5 4.5 35	16 35.5 16 38.7 17 49 20 40.9	6 11 7 50 15	
12 12 12 13	2 Simple 2 3 Simple 3 A 6 Complex 1 Simple 1 6 Complex	12 38 14 33 14 34.3 22 32 14 46	4 30 4 3 5	12 39.5 indet. 14 35 22 33.3 14 48.6	11 11 16 7 27	
13 13 14	2 Simple 2 6 Complex f 6 Complex	17 12.8 23 21.5 11 19.5	3 18 4	17 14 23 22.9 11 20.5	16 34 27	,
14	4 Post Increase 1 Simple 1 A 2 Simple 2	14 07.5 14 07.7	25 3 0.3	14 08.5 14 07.8	10 7 22	
14 14 14	1 Simple 1 3 Simple 3 8 Group (2)	15 15 17 06 21 15.8	2.5 20 5.2	15 16 17 09	7 7	
14	2 Simple 2 2 Simple 2	21 15.8	3	21 17.2 21 20.7	70 8	
15 15 15	3 Simple 3 A 6 Complex 3 Simple 3 A 1 Simple 1	13 49 18 25 20 45.5 23 02 23 14.5	1 10 3 45 8 1	14 00 18 48 20 47.8 indet. 23 15	8 9 34 5 7	
18 18 19 19	2 Simple 2 2 Simple 2 2 Simple 2 2 Simple 2	18 29.3 23 47.5 9 42.8 11 27	3•5 2•5 5 3	18 30.2 23 48 9 44.2 11 28	43 9 140 10	In interference

#### SOLAR RADIO EMISSION OUTSTANDING OCCURRENCES JUNE 1958

OTTAWA

June	Туре*	Start UT	Dumand	<del></del>		2800 M
1958	Type	Hrs:Mins	Duration Hrs:Mins	Maxim Time UT Hrs:Mins	Peak Flux	Remarks
19 19 19	8 Group (2) 1 Simple 1 1 Simple 1 2 Simple 2 6 Complex 4 Post Increase f	12 57.1 12 57.1 13 02.5 13 30 14 37.9	7.9 1.5 2.5 3 4	12 57.5 13 03.4 13 31 14 39.5	7 6 23 48 10	
19 19 19 20	1 Simple 1 1 Simple 1 1 Simple 1 8 Group (2) 1 Simple 1 1 Simple 1	18 43.5 20 34 21 29 10 50 10 50 10 53	5 2.5 3 4 2 1	18 45 20 35 21 30.5 10 50.8 10 53.5	6 3 4 6 5	
20 21 22 23	2 Simple 2 3 Simple 3 A 1 Simple 1 3 Simple 3 f 1 Simple 1	17 59.5 12 35 12 35 17 55 13 14	6 1 8 1 40 5	18 02 indet. 12 38 18 15 13 16	12 6 7 12	
25 26 26 27	2 Simple 2 f 1 Simple 1 1 Simple 1 2 Simple 2	16 26 16 45 22 42 21 43.3	17 4 3 1	16 27.4 16 46.5 22 44 21 43.6	200 4 6 <b>1</b> 5	
28	3 Simple 3 A 6 Complex 8 Group (2) 2 Simple 2 2 Simple 2	15 00 17 02.5 18 39 18 39 18 43	5 20 1.5 9 2 5	17 45 17 02.7 18 39.4 18 45.2	23 5 12 8	
29 29 29 29	2 Simple 2 2 Simple 2 1 Simple 1 1 Simple 1	9 36.3 13 18.5 15 09.5 18 02.5	1 4 0.5 2	9 36.8 13 20.5 15 09.7 18 03.2	35 8 4 5	In sunrise osc.
29 29	2 Simple 2 f 4 Post Increase 8 Group (2) 1 Simple 1 1 Simple 1	20 25 21 29 21 29 21 35	3.5 7 7 1 1	20 25.8 21 29.5 21 35.5	145 10 7 3	
	1 Simple 1	21 35			3	

COMMERCE - STANDARDS - BOULDER

#### HOURS OF OBSERVATIONS: APRIL, MAY, JUNE 1958

OBSERVING PERIOD: April 1055 UT - 2315 UT (approx.)
May 1005 UT - 2355 UT (approx.)
June 1000 UT - 2400 UT (approx.)

#### with the following exceptions:

- (1) No observations:
  - April 18 1405-1545
- (2) Observations commenced: May 7 at 1140

June 24 at 1035

25 1035

1035 26

30 1035 (3) Periods of interference obscuring the records on: April 3, 5, 11, 15, 16, 24.

May 3, 5, 6, 12, 13, 14, 15, 16, 17, 19, 20, 22.

June 20, 23, 24, 28.

#### SOLAR RADIO EMISSION

### DAILY DATA MAY 1958

CORNELL

200 MC

COMME				T			
May 1958	10-22 <sub>w</sub> Ho	Density m <sup>-2</sup> (c/s urs UT 15 18	y )-1 18 21	0	bility to 3 rs UT 15 18	18 21	Observing Periods Hours UT
	15	10	51	15	10	21	
1 2 3 4 5	[[ 22 32 [ 18] [ 14 [ 15	21 32 15]] 13]] 33	22] 20] 69]]	[[1 2 [1] [1 [1	2 1 1]] 1]] 2	2] 1] 2]]	1420-2000 1235-2005 1255-1415,1455-1600 1255-1330,1350-1610 1245-1925
6	[ 13	15	15]	ĽΟ	1	1]	1245-2000
7 8 9	[ 12	13	12]]	[1	1	1]]	1240-1830
10	[ 12	11]]		[0	0]]		1245-1610
11 12 13 14 15	11 [11 [11 [12 [11	11]] 11 12 11 11	11]] 12]] 11] 11]	0 [1 [0 [0	0]] 1 1 0 0	0]] 1]] 0] 0]	1205-1600 1250-1850 1245-1425,1455-200 1255-1835,1845-2010 1250-1520,1555-2015
16 17 18 19 20	11 17 [ 11 11 [ 11	11 14]] 11]] 11 11	11] 11] 11]	0 2 [1 [0 [0	0 2]] 0]] 0 0	0	1240-2010 1235-1600 1250-1605 1240-2010 1240-2000
21 22 23 24 25	CC 12 12 E 12 12 C 12	12 12 12 12]] 12]]	12] 12] 12]	[[0 0 [1 0 [1	0 0 0 0]]	0] 0] 1]	1415-2010 1235-2020 1325-2020 1220-1600 1250-1600
26 27 28 29 30 31	[ 12 [ 12 12 12 [ 12 12	12 12 12 12 12 12	12] 12]] 12] 12] 12]	[0 [0 0 0 [1 0	0 1 1 0 1 0]]	0] 1]] 1] 1] 0]	1245-2020 1250-1835 1240-2010 1240-1935 1245-2010 1230-1600

<sup>[ = 1</sup>st hour missing.

<sup>[[ = 1</sup>st two hours missing.

<sup>] =</sup> last hour missing.

<sup>]] =</sup> last two hours missing.

## SOLAR RADIO EMISSION OUTSTANDING OCCURRENCES MAY 1958

CORNELL

200 MC

May	Type	Start	Time of	Duration	Type	_10 <sup>-22</sup> w m	x Density -2(c/s)-1	
1958	Ap.J	UT	Maximum	Minutes	IAU	Inst.	Smooth	Remarks
1 2 3	3 7,4 3	1750.5* 1335 1538	1751* 1539*	1.5* 183 1.5	SA* F CA	630* 72*	450* 41*	
5	3 7	1534 1703.5		.5 > 160	CA F	180*	120*	
6 8	2 3 3	1605 1927.5 1507.5	1928	33 3	F CA	73 880*	47 (204)	
10 12	3 3	1307.3 1459 1341		.5 1 .5	CD CD SD	880* 44 550*	630* 23 450*	
13	2 8 3	1545 1618 1809.5	1548.5* 1620*	6.5 4.5 .5	E CD CD	91* 13,000* 400*	72* 11,000* 280*	
15 17	3 9	1301 1349.5	1351*	< .25 7	CD ECD	91* 320*	72* 210*	
19	9 8 3	1435 1825 1851	1829*	52 4.5 .5	F CD CD	210* n65	180*	off-scale on linear record
	3	1854		.5	CD	n65		off-scale on linear record
21	3	1647		1	CD	45	27	
23	3 2 3	1339 1437 1725.5	1445.5*	< .25 14 1	CD F CD	260* 450* n65	210* 320*	
	2 3	1842 1940	1844*	11 1	F CD	91* n65	72*	
25 27	2 3 3	1346.5 1537.5 1651.5	1349*	3.5 .5 .5	CD CD	140* 120* 36	95* 72* 23	
28	3 3	1506.5 1641.5 1823.5 1832	1642.5 1824.5*	9 1.5 1.5	CD CA CA CA	n65 n65 120* n65	91*	
29	3 2 3 2	1905 1932 1816	1832.5 1905.5 1932* 1819*	1 3 1 9	CA CA CA	n65 n65 37 3600*	22 2800*	
30	2 3 8	1338 1442.5 1510	1340.5* 1513.5*	4.5 .5 10	CD CD CD	29 91*	16 72*	

<sup>\* =</sup> Logarithmetic Recorder

#### SOLAR RADIO EMISSION

### DAILY DATA JUNE 1958

CORNELI.

200 MC

CORNEL							200 1112
June 1958	10-22w	x Densit m <sup>-2</sup> (c/s ours UT 15 18	y )-1 18 21	0	to 3 irs UT 15	18 21	Observing Periods Hours UT
1 2 3 4 5	12 15 13 12 [[14	12]] 15 13 13 14	16] 13] 13] 15]	0 1 1 0 [[1	133 1 1 1 1	13 13 13 13	1220-1600 1230-2005 1235-2000 1230-2010 1315-2010
6 7 8 9 10	16 31 [16 15] 12	17 22]] 15]] 16 13	24J 13J 13	2 2 [1 1] 1	2 2]] 1]] 0 1	3] 1) 1	1230-1330,1335-2000 1235-1605 1255-1615 1240-1415,1500-2005 1235-2035
11 12 13 14 15	12 [[12 [12	11 12 [12 12]] 12]]	12] 12]	0 [[0 [0	0 0 [0 1]] 0]]	0]	1240-1900 1240-2005 1550-1940 1230-1600 1245-1610
16 17 18 19 20	12 12 £12 £12 £12	12 12 12 11	12] 12] 12] 11] 11]	0 1 [0 [1 [0	0 0 0 1 0	0] 0] 1] 1] 0]	1240-2005 1235-2010 1310-2020 1240-2000 1245-2000
21 22 23 24 25	11 12 [11 [11 [18	11]] 12]] 11 11 15	11]	0 1 [0 [0 [2	0]] 0]] 0 0 0 2	OJ	1230-1600 1240-1600 1240-1810 1245-2000 1240-1805
26 27 28 29 30	19 12 11 11 11	15 12 12]] 11]] 11	113	2 0 1 1 0	1 1 2]] 0]] 1	1]	1240-1800 1235-1800 1230-1600 1235-1600 1235-2005

<sup>[=</sup> lst hour missing.

<sup>[[ = 1</sup>st two hours missing.

<sup>]=</sup> last hour missing.

<sup>]] =</sup> last two hours missing.

## SOLAR RADIO EMISSION OUTSTANDING OCCURRENCES JUNE 1958

CORNELL

200 MC

June	Type	Start	Time of	Duration	Type	Max. Flux Density 10-22 <sub>w m</sub> -2(c/s)-1		200 140
1958	Ap.J	UT	Maximum	Minutes	IAU	Inst.	Smooth	Remarks
2	2 3	1418 1548.5	1419 1550	1.5 2 1	CD CD	530 91	450 72	
3	3 3 3	1850 1339 1814	1850.5	< .25 .5	CD SD CD	~ 65 120 ~ 65	91	
5	2 2 8	1721 1810.5 1617	1810.5	13.5 4.5 18	CD CD ECD	~ 65 ~ 65 ~ 65		off-scale on linear
	8	1702		33	ECD	~ 65		record-1622.5-24.5 UT off-scale on linear record-1708.5-23 UT
6	6	ъ1228		> 452	F			
7	3	1556.5		1.5	CA	~ 65		
8	2	1451		4	CD	<b>∼</b> 65		
14	2	1518.5		6	CD			
17	3	1647.5		.25	CD	48	32	
18	2	1923		2	CD	48	34	
19	3	1703.5		< .25	CD	72	41	
20	3	1517		1	CD	48	35	
22	1	1242		108	M			
25	2	1410		82	F			
28	2	1516.5	1517.5	5	F	60	36	
29 29	2	1237.5		15	F	92	66	
30	3	1251.5 1653	1653.5	.5 2	CD CD	82 30	66 15	

## SOLAR RADIO EMISSION DAILY DATA APRIL 1958

BOULDER

167 MC

ROOF	BOULDER 167 MC								
		Flux Density $10^{-22} \text{w m}^{-2}(\text{c/s})^{-1}$ Variability 0 to 3			Variability Observing Periods O to 3				
			Hours l	JT			Hours UT Hours UT		
Apr. 1958	0 3	12 15	15 18	18 21	21 24	Day			
1 2 3 4 5	- - - -	-	47 28 21 24 22	57 3 <sup>4</sup> 25 18 20	54 142 18 16 18	52 64 21 21 21	3 3 2 3 14.5-18.0,18.3-01.2 - 1 2 2S 2 2 12.8-01.2 - 2 2 2S 2 2 12.8-23.4		
6 7 8 9	- - - -	-	19 21 18 16 18	26 17 18 20 17	26 17 23 18 17	23 19 20 18 17	- 1 2 1S 2 2 12.7-01.2 - 2 1 2S 2 2 12.6-01.3 - 2 1 2 2 2 13.7-01.3		
11 12 13 14 15		-	16 - 17 16 17	15 17 17 18 18	16 16 16 17 17	16 17 17 17 17	0S 1S 1 1S 17.8-01.3 - 2 2 2 1S 2 12.5-01.3 - 1 1 2 1S 1 12.5-15.5,16.3-01.3		
16 17 18 19 20	- - -	-	17 21 19 18 17	17 20 21 18 17	16 18 16 17 16	17 20 19 18 17	- 2 18 28 28 28 12.4-01.3 - 2 2 18 18 18 12.4-01.3 - 1 1 28 28 28 12.4-01.4 - 08 1 18 18 18 12.3-01.4 - 1 1 1 2 1 12.3-15.8,16.5-01.4		
21 22 23 24 25	- - - -	-	16 19 18 17 19	16 18 - 16 19	14 - - 13 17	15 20 17 16 18	- 1S 1 1 1S 1S 12.3-13.5,14.3-01.4 - 2S 2 1 1 1 12.3-01.4 - 1 1S 1S 0S 1S 12.3-01.4 - 0S 1 0S 0S 14.3-23.5 - 0S 2 2S 1S 1S 13.8-01.5		
26 27 28 29 30 31	1 1 1 1	-	18 21 21 37 3 <sup>4</sup>	17 21 22 47 35	15 22 24 45 33	17 21 22 41 35			

## SOLAR RADIO EMISSION OUTSTANDING OCCURRENCES JUNE 1958

CORNELL

200 MC

CORNEL	L							200 MC
June 1958	Type Ap.J	Start UT	Time of Maximum	Duration Minutes	Type I AU	Max. Flux Density 10-22w m-2(c/s)-1 Inst. Smooth		Remarks
1730	1.7.5		11071211011	TELEGO CO			DAIOUGII	Кещатка
	١.							
2	2	1418	1419	1.5	CD	530	450	
	3	1548.5	1550	2	CD	91	72	
	3	1850	1850.5	1	CD	~ 65		
3	3	1339		< .25	SD	120	91	
	3	1814		.5	CD	<b>~</b> 65		
4	2	1721		13.5	CD	~ 65		
4	2	1810.5	1810.5	4.5	CD	~ 65		
5	8	1617	1010.5	18	ECD	~ 65		off-scale on linear
,	ľ	1017		10	ECD	~ 03		record-1622.5-24.5 UT
	8	1702		33	ECD	~ 65		off-scale on linear
	ľ	1702		33	202	03		record-1708.5-23 UT
	1	1						100010 1700.5 25 01
6	6	ь1228		> 452	F	1		
7	3	1556.5		1.5	CA	<b>~</b> 65		
8	2	1451		4	CD	~ 65		
14	2	1518.5		6	CD			
17	3	1647.5		.25	CD	48	32	
	1	l						
18	2	1923		2	CD	48	34	
19	3	1703.5		< .25	CD	72	41	
20	3	1517		1	CD	48	35	
22	1	1242		108	м			
25	2	1410		82	F			
		1					_	
28	2	1516.5	1517.5	5	F	60	36	
29	2	1237.5		15	F			
29	3	1251.5		.5	CD	82	66	
30	3	1653	1653.5	2	CD	30	15	

## SOLAR RADIO EMISSION DAILY DATA APRIL 1958

BOULDER

167 MC

BUUL	BOULDER 167 MC												
			Flux O-22 <sub>w m</sub>	Densit	y \_1					abil			Observing Periods
		1	O_55_W U	(c/s	3) -				(	) to	3		
			Hours U	T		1		Но	urs	UT			Hours UT
Apr.	0	12	15	18	21	Day	0	12	15	18	21	Day	
1958	3	15	18	21	24		3	15	18	21	24		
1		_	47	57	54	52	_	2	2	2	2	2	13.5-01.1
2	_	_	28	34	142	64	_	-	3	3	2	3	14.5-18.0,18.3-01.2
3	_	-	21	25	18	21	_	1	2	2S	2	2	12.8-01.2
4	-	-	24	18	16	21	-	2	2	25	2	2	12.8-23.4
5	-	-	22	20	18	21	-	1	1	2	1	1	12.7-01.2
6	_	_	19	26	26	23	_	2	2	2	2	2	12.7-01.2
7	_	_	21	17	17	19	-	1	2	1S	2	2	12.7-01.2
8	-	-	18	18	23	20	-	2	1	25	2	2	12.6-01.3
9	-	-	16	20	18	18	-	2	1	2	2	2	13.7-01.3
10	-	-	18	17	17	17	-	0	2	2S	2S	2S	13.8-01.3
11	_	_	16	15	16	16	_	-	18	25	2S	25	15.8-01.3
12	-	-	-	17	16	17	_	-	OS	18	1	18	17.8-01.3
13	-	-	17	17	16	17	-	2	2	2	18	2	12.5-01.3
14 15	-	-	16 17	18 18	17 17	17	- -	1	1 1	2 1S	1S 1S	1 1S	12.5-15.5,16.3-01.3 12.4-20.6,21.8-01.3
1 19	-	-	Τ (	10	Τ (	17	-	Τ.	Τ.	Тр	TD	l lo	12.4-20.0,21.0-01.3
16	-	-	17	17	16	17	-	2	1S	28	2S	28	12.4-01.3
17	-	-	21	20	18	20	_	2	2	18	1S	1S	12.4-01.3
18	-	-	19	21	16	19	-	1	1	2S	2S	2S	12.4-01.4
19	-	-	18 17	18 17	17 16	18 17	-	OS 1	1 1	1S 1	1S 2	lS l	12.3-01.4 12.3-15.8,16.5-01.4
20	_	-	<b>-</b> [	+1	10	7.	_	_	_	_	_	+	12.3-17.0,10.7-01.4
21	-	-	16	16	14	15	_	18	1	1	1S	18	12.3-13.5,14.3-01.4
22	-	-	19	18	-	20	-	2S	2	1	1	1	12.3-01.4
23	-	-	18	-	-	17	-	1	1S	lS	OS	1S	12.3-01.4
24 25	_	-	17 19	16 19	13 17	16 18	-	- 0S	0S 2	1 2S	OS 1S	OS 1S	14.3-23.5 13.8-01.5
-	-	-	⊥フ	エフ	Τ1	10	-	CD	۷	دے	TD	10	13.0-01.)
26	-	_	18	17	15	17	_	1	1	1	lS	1	12.2-01.5
27	-	-	21	21	22	21	-	1	1S	1S	1S	18	12.2-01.5
28 29	-	-	21 37	22 47	24 45	22 41	-	1 2	2 2S	2S 2S	2S 1S	2S 2S	12.2-01.6 12.1-01.6
30	-	-	31 34	47 35	45 33	35	-	2 1S	25	25	25	25	13.8-01.6
31	_	_	٠, ر	57	رر	ارد		TD	_	_	_	_	15.0 01.0

## SOLAR RADIO EMISSION SPECTRUM OBSERVATIONS JUNE 1958

FORT DAVIS

100-580 Mc

DATE and OBSERVING TIMES (U.T.) 1958	TYPE I (NOISE STORMS and CONTINUUM)	TYPE II (SLOW DRIFT BURSTS) & UNCLASSIFIED	TYPE III (FAST DRIFT BURSTS)	REMARKS
June 1 0000-0145 1233-2400		Uncl. 0024 3 Uncl. b 1745 1 Uncl. b 1757 1 Uncl. g 1808 1- Uncl. b 1819 1- Uncl. b 1827 1 Uncl. b 2132 2 Uncl. g 2325 1 Uncl. g 2330-31 2	b 0022 1 g 0024 1 b 1331 1- g 1730 2 g 1743 2 b 1934 1- b 1941 3 g 1948-50 1 g 1951 2 g 2126-27 2 b 2135 2 g 2251-52 2	
June 2 0000-0145 1232-2400		Uncl. g 0021 1 Uncl. g 1550 1 Uncl. g 1704 1 Uncl. b 2252 1-	g 0042-0046 1 b 0135 g 1418-19 2 g 1549-50 1 b 2251 1	
June 3 0000-0144 1218-2400		Uncl. b 1815 1-		
June 4 0000-0145 1218-2400	Cont. IV 0047-48 1 Cont. IV 2037 2 Cont. IV 2142-43 1- Cont. IV 2143-45 2 Cont. IV 2148-53 3 Cont. IV 2153-59 2 Cont. IV 2159-2203 3 Cont. IV 2203-2205 2 Cont. IV 2205-2209 1	Uncl. g 1324 1- Uncl. g 1835 1- Uncl. g 1842 1 Uncl. 2149 3 Uncl. b 2152 3	G 1811-12 1 b 2144 2 G 2147-51 2 b 2152 3	
June 5 0000-0145 1218-2400		Uncl. 1621.2-29 3	b 0101 1- g 1256 2 g 1357 1 g 2127-28 2	
June 6 0000-0145 1218-2400	1222 1	Uncl. g 0011 1 Uncl. g 0051-53 2 Uncl. b 1707 3	g 0130-0132 2	
June 7 0000-0004 0006-0145 1220-2400	0004 1- 0007-0139 1 1220-1432 1 1503-1802 1- 2034 1- 2311 1- 2346-47 1	Uncl. 1816 2		
June 8 0000-0145 1218-2400	2345 1-	Uncl. 1453-54 1- Uncl. 1455 1- Uncl. b 1603 1-	g 1232 2 g 1452-53 3 g 1649-50 3 g 2041 2 b 2104 3 b 2131 2 g 2253-55 3	
June 9 0000-0145 1220-2400	2102-03 1- 2301 1- 2349 1-	Uncl. 2135 1-	ъ 1311 3	
June 10 0000-0150 1219-2400	0046 1- 1845 1- 2009-14 1 2151 1- 2356 2	Uncl. g 1224 1-	b 1543 1-	

## SOLAR RADIO EMISSION SPECTRUM OBSERVATIONS JUNE 1958

FORT DAVIS 100-580 Mc

DATE and OBSERVING TIMES (U.T.) 1958  TYPE I (NOISE STORMS and CONTINUUM)					TYPE II (SLOW DRIFT BURSTS) & UNCLASSIFIED			PE III (FAS		REMARKS
June 11 0000-0148 1219-1538 1551-2400		2101 <b>-2</b> 114 2231 <b>-</b> 44		Uncl.	0135-38	1-				
June 12 0000-0150 1218-1600 1800-2400				Uncl. g Uncl. b Uncl. g Uncl. b	1227	1- 1- 1 2	ପ୍ର ଓ ଓ ଓ	1220 1225 1327 1333 1552	3 1 1- 2 3	
June 13 0000-0147 1220-2400	Cont.	1449	2				g	1448-49	2	
June 14 0000-0150 1219-2400	Cont.	1520	3	Uncl. Uncl. b Uncl.	1709	3 1- 1- 3	6 8 8 8 9 9 9 9 9 9 9 9 9 9 9 9 9 9 9 9	1422 1431 1518 1520 1522-23 1525 1709 1741-42 1743-44 2121 2205	2 2 2 3 3 1- 3 2 1	
June 15 0000-0150 1219-2400				Uncl.	2047 2048	2 3	8 b b 8 b	0012 1226 1931 2042 2046 2048	3 2 1- 1- 1- 3	
June 16 0000-0150 1219-2400							8 6 8 6	0026 1258 1300 1605 1807-08	3 1- 2 1- 3	
June 17 0000-0150 1218-2400				Uncl. b Uncl. b Uncl. b Uncl. b	1648 1842	1- 1- 1- 1-	b b b	1349 1420 1943 2202-03 2206	1- 1 3 1- 1-	
June 18 0000-0150 1218-2400				Uncl. b Uncl. g Uncl. g	1719	1- 1- 1-	ъ 8 в ъ ъ	0055 1225 1341 1343 2005 2039	1 1 2 2 3 2	
June 19 0000-0149 1218-2400	Cont.	0131 1307 1954	2 1 1				80 80	0130-31 1331	1	1331 Inverted U burst.
June 20 0000-0150 1218-2400		0050	1	Uncl. b Uncl. b Uncl. b Uncl.	2317	1- 1- 1- 1-	ь	1518	3	
June 21 0000-0130 1444-2400	Cont.	2320-21	3	Uncl. g Uncl. g Uncl. b Uncl.	2249	1 1 1 1-	b b g b	1805 1809 2204-05 2321	1- 1- 3 1	

### SOLAR RADIO EMISSION SPECTRUM OBSERVATIONS

JUNE 1958

FORT DAVIS

100-580 Mc

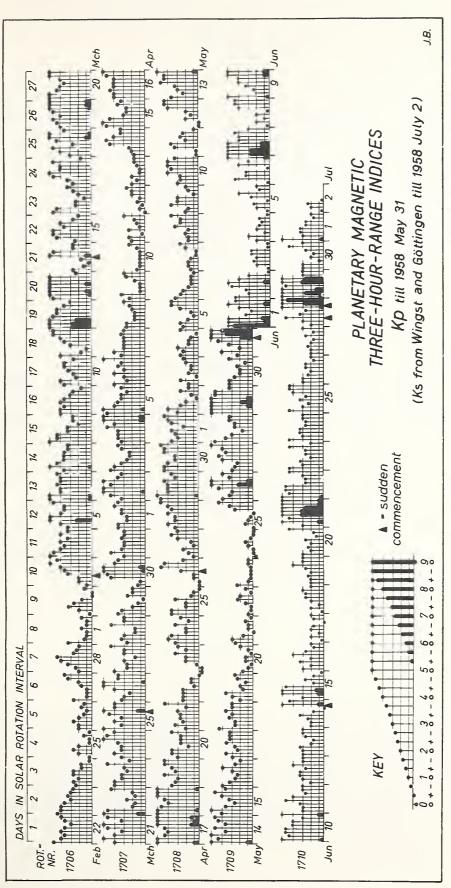
FORT DAVIS				100-580 Mc
DATE and OBSERVING TIMES (U.T.) 1958	TYPE I (NOISE STORMS and CONTINUUM)	TYPE II (SLOW DRIFT BURSTS) & UNCLASSIFIED	TYPE III (FAST DRIFT BURSTS)	REMARKS
June 22 0000-0150 1219-2144 2145-2400	1238-1301 1-	Uncl. b 1555 2	g 1221 1 b 1543 1	
June 23 0000-0149 1219-2400		Uncl. b 0029 2 Uncl. g 0053 1		
June 24 0000-0150 1219-2400	2154-2335 1-	Uncl. b 0007 2		
June 25 0000-0149 1220-2400	0003-04 1- 0041-54 1- 0132-35 1- 1243-1446 1 1446-1449 2 1449-1737 1 1737-1804 2 1804-2007 1 2007-2023 2	Uncl. g 2227 2	b 2133 1-	
	2023-28 3 2028-47 2 2 2034 3 2047-2108 3 2108-23 2 2152-2204 1 2248-2334 1 2334 2			
June 26 0000-0150 1219-2400	0143 2 1221-49 1 1357 1 1425-49 1- 1449-1506 2 1710 1- 1807 1- 1939-49 1- 2010 1	Uncl. b 1825 1	b 2020 1-	
June 27 0000-0150 1220-2400	0146 1 1623-1719 1- 2132 1- 2202-2347 1-	Uncl. g 1818 3	b 0006 3 b 1818 3 g 2055 3 g 2143 3	
June 28 0000-0149 1220-2325 2326-2400	1755-1913 1 1942-2020 1- 2042-2313 1 2313-2325 2 2326-41 2 2341-53 1 2353 2	Uncl. 1703 1-	b 1659 1- b 1744 1- g 1845-46 3	
June 29 0000-0149 1219-2400	0017 2 0017-0143 1 Cont. 2025 2	Uncl. 1733 1	8 0029 3 b 0133 1- 8 1510 3 8 1750 1 8 1751-52 2 b 1754 2 8 2016 2 8 2024-25 3 8 2127 1 b 2129 1	2016 Inverted U burst.
June 30 0000-0150 1213-2400		Uncl. g 1732 1	b 0014 1- g 1227 2 b 1237 1- b 1436 1 b 1451 3 g 1654-55 1- b 2054 3	



### GEOMAGNETIC ACTIVITY INDICES

May 1958

May 1958	С	Values Kr Three hour Gr. 1 2 3 4		Sum	Ар	Final Selected Days
1 2 3 4 5	1.0 0.7 0.2 0.3 0.7	30 40 3- 20 1 2- 3- 1- 2- 1 3- 2+ 30 2- 2	2+ 4- 40 4+ 1- 1+ 3- 2- 1+ 1+ 20 2+ 2- 1+ 1- 3- 1+ 3- 4- 4-	280 180 14- 160 190	21 11 7 8 12	Five Quiet 7 20
6 7 8 9 10	0.3 0.2 0.8 0.7	1+ 2- 2- 1+ 1 2- 2- 20 2- 3 20 2- 4- 30 3	+ 1+ 1+ 10 - 1- 20 3- + 3- 3- 4- 0 10 2- 20 - 40 40 3+	150 120 19+ 180 27+	8 6 11 10 20	22 23 24
11 12 13 14 15	0.5 0.8 1.3 1.4	0+ 0+ 20 1+ 2 4- 2+ 30 3+ 5 6- 4+ 50 4- 5	0 3- 2+ 10 0 40 4+ 4+ 1- 4+ 4+ 6- 1- 4- 4+ 5- 1- 40 5- 4-	16+ 19- 31+ 360 310	10 14 29 38 27	Five Disturbed 13 14
16 17 18 19 20	0.9 0.9 1.0 0.6 0.2	30 4- 2+ 3- 4 4+ 3- 3- 3+ 3 20 30 4- 2- 2	o 4o 3- 2o - 3+ 3+ 3+ o 4- 3o 3o o 1+ 2+ 3- + 1- 1o 2o	240 25+ 26- 19- 10-	16 17 18 10 5	26 29 31
21 22 23 24 25	0.2 0.1 0.1 0.1 0.6	10 1- 1- 1+ 2 2+ 10 1- 1- 1 00 0+ 0+ 10 1	o 1o 1+ 1- + 2o 1- 1+ - 1- 1o 1o + 1- 0+ 0+ + 3- 3+ 4+	12- 100 80 4+ 140	6 5 4 2 10	Ten Quiet 3 4
26 27 28 29 30 31	1.3 1.1 1.0 1.6 0.8 1.7	30 3- 2+ 40 4 5- 4- 30 2- 3 50 50 6- 6+ 6 3+ 3+ 3- 20 2	- 6- 3+ 40 + 4- 40 50 - 3- 3+ 3+ - 5- 5- 3+ 0 2+ 1+ 3+ + 6+ 8- 80	34+ 290 250 40+ 20+ 380	39 24 18 52 12 72	6 7 20 21 22 23 <b>24</b>
Mean:	0.75			Mean:	17	25



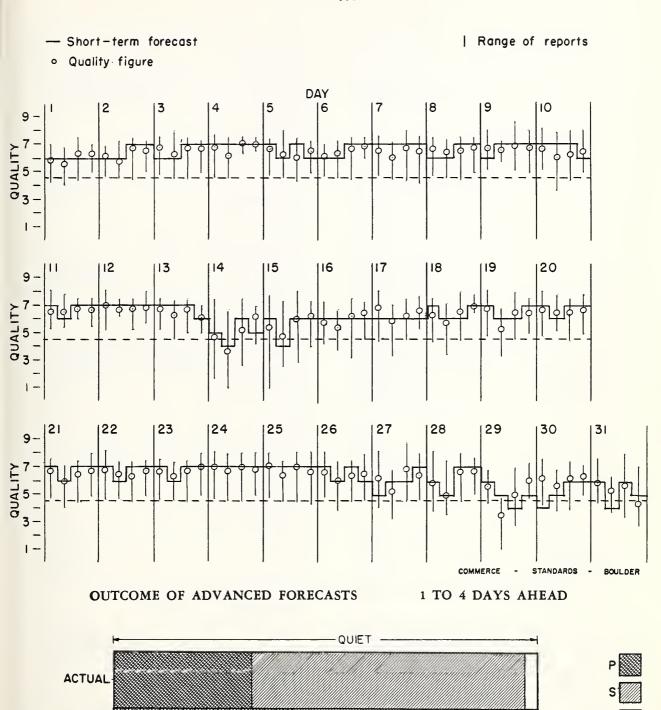
OMMERCE - STANDARDS - BOULDER

## CRPL RADIO PROPAGATION QUALITY FIGURES AND FORECASTS NORTH ATLANTIC

MAY 1958

May 1958	North Atlantic 6-hourly quality figures	Short-term forecasts issued about one hour in advance of:	Whole Advance forecasts Geomag- day (J-reports) for netic index whole day; issued in advance by:
	00 06 12 18 to to to to 06 12 18 24	00 06 12 18	1-4 4-7 8-25 Half Day days days days (1) (2)
1	60 6- 6+ 6+	6 6 6 6	60     7     7       6+     6     7       7-     6     7       2     2       6+     7     2       2     2       2     2       3     1       2     2       2     2       3     1       2     2       3     2       2     2       3     3       4     4       5     7       2     3
2	6+ 60 7- 7-	6 6 7 7	
3	7- 6+ 7- 7-	6 6 7 7	
4	70 6+ 70 70	7 7 7 7	
5	7- 6+ 60 7-	7 6 7 6	
6	6+ 6+ 7- 7o	6 6 7 7	7- 6 5 2 2
7	7- 6+ 7o 7-	7 7 7 7	7- 6 6 6 2 2 7
8	7- 7- 7- 7-	6 6 7 7	7- 6 6 6 2 (4)
9	7- 7- 7o 7-	6 7 7 7	7- 7 6 3 2
10	7- 6o 6+ 7-	7 7 7 6	6+ 7 6 3 (4)
11	7- 7- 7- 7-	7 6 7 7	7- 7 6 2 2
12	70 7- 7- 70	7 7 7 7	70 7 6 1 3
13	7- 6+ 7- 60	7 7 7 6	6+ 7 7 3 (4)
14	5- 4- 5+ 6+	5 4 6 5	5- 7 7 (4) 3
15	5+ 5- 60 6+	6 4 6 6	6- 7 7 (4) 3
16	6- 5+ 6+ 7-	6 6 6 6	60     6     7     3     2       6+     6     7     3     3       6+     7     7     3     3       6+     7     7     3     2       7-     6     7     1     2
17	70 60 6+ 7-	6 6 6 6	
18	6+ 6- 7- 70	7 6 6 7	
19	70 5+ 7- 7-	7 6 6 7	
20	7- 7- 7-	7 6 7 7	
21 22 23 24 25	7- 60 7- 7- 70 7- 6+ 7- 7- 6+ 7- 70 70 7- 70 7- 70 7- 70 7-	7 6 7 7 7 6 7 7 7 6 7 7 7 7 7 7 7 7 7 7	7- 6 7 2 2 7- 6 7 1 2 7- 7 7 1 1 70 7 7 7 2 3
26	7- 60 6+ 7-	7 6 7 6	6+ 7 7 7 (4) (4) 6+ 5 7 3 3 60 6 7 3 3 3 65) (5) (4) (4) 6+ 5 7 3 2 3 (5)
27	6+ 5+ 70 7-	5 6 6 7	
28	60 50 7- 7-	6 5 7 7	
29	6- 3+ 50 60	6 5 4 5	
30	6+ 6- 6+ 6+	4 5 6 6	
31	60 5+ 6- 4+	6 4 6 5	
Score	·	P 23 15 22 23 S 7 14 9 7 U 0 0 0 0 F 1 0 0 0	10 9 20 18 1 4 0 0
		P 0 1 0 0 S 0 0 0 1 U 0 1 0 0 F 0 0 0 0 d values.	0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0

## CRPL RADIO PROPAGATION QUALITY FIGURES AND FORECASTS NORTH ATLANTIC MAY 1958



COMPARISON (SEE TEXT)

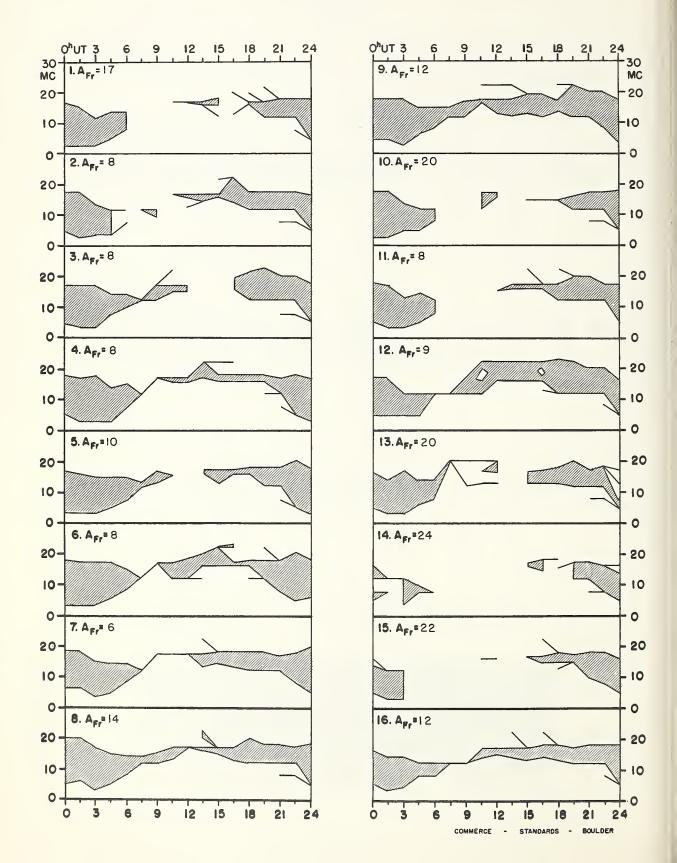
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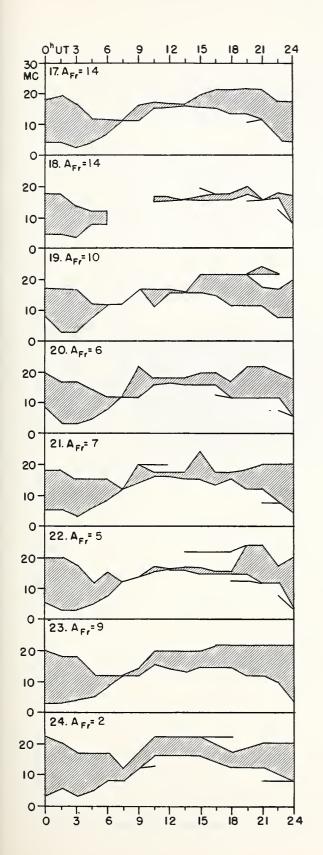
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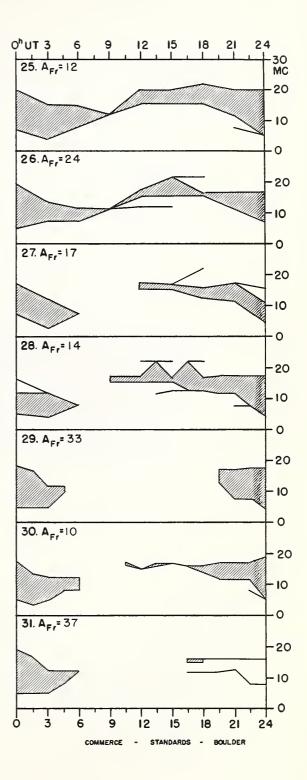
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31

## USEFUL FREQUENCY RANGES -- NORTH ATLANTIC PATH MAY 1958







## CRPL RADIO PROPAGATION QUALITY FIGURES AND FORECASTS NORTH PACIFIC MAY 1958

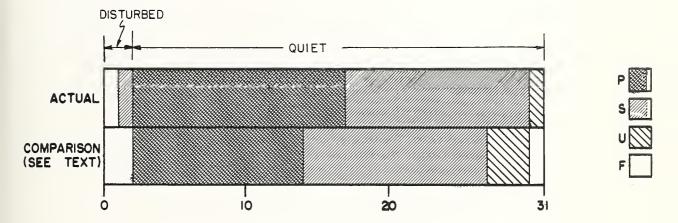
May 1958	North Pacific 8-hourly quality figures	Short-term fore- casts issued at	Whole day index	Advance forecasts (Jp reports) for whole day; issued in advance by:	Geomag- netic K <sub>Si</sub>
	03 11 19 to to to 11 19 03	02 10 18		1-4 4-7 8-25 days days days	Half Day (1) (2)
1 2 3 4 5	4 6 6 5 6 6 4 5 7 5 6 6 4 6 6	6 6 6 5 5 6 6 6 6 5 6 6 5 5 6	5 6 5 5	4 7 5 7 5 5 4 5 4 5	3 2 3 1 2 1 2 1 2 2
6 7 8 9 10	6 6 7 6 6 6 5 6 6 6 7 6 6 6 5	5 6 6 6 6 6 6 6 6 6 6 6	7 6 6 7 6	5 5 6 6 6 6 6 6 7 6	2 1 1 2 2 2 2 2 3 3
11 12 13 14 15	6 6 6 6 6 6 5 5 6 4 4 5 5 5 6	6 6 6 5 6 6 5 6 6 5 4 5 4 5 6	6 6 6 (4) 5	7 6 7 7 7 7 5 7 6 7	2 2 1 (4) (4) (4) (6) (4) (4) (4)
16 17 18 19 20	5 6 6 6 6 6 5 5 6 6 6 6 6 5 6	5 5 6 6 6 6 6 6 6 6 7 7	6 6 6 6	6 7 6 6 6 6 6 6 6 6	3 3 3 3 (4) 3 3 2 1 1
21 22 23 24 25	6 6 6 6 6 7 6 6 7 6 6 7 7 6 8	6 6 6 6 7 6 7 7 7 7 6 7 6 7 7	6 6 7 7 7	6 6 7 6 7 6 7 7 6 6	2 1 1 2 1 1 0 1 1 2
26 27 28 29 30 31	6 4 5 6 5 6 6 6 6 4 2 5 6 6 6 6 5	7 6 5 5 5 6 5 5 6 5 3 4 5 5 6 6 6 6	6 6 (4) 6 6	6 6 6 6 6 7 6 7 7 7 6 7	(4) (5) (4) (4) (4) 3 (6) (4) (4) 3 (4) (6)
Score:	Quiet Periods	P 14 15 23 S 12 12 8 U 0 1 0 F 0 0 0		15 16 13 10 1 3 0 0	
1	Disturbed Periods	P 0 1 0 S 3 1 0 U 0 0 0 F 2 1 0		0 0 1 0 0 0 1 2	

<sup>( )</sup> represent disturbed values.

### CRPL RADIO PROPAGATION QUALITY FIGURES AND FORECASTS NORTH PACIFIC

MAY 1958

OUTCOME OF ADVANCED FORECASTS 1 TO 4 DAYS AHEAD



### ALERT PERIODS AND SPECIAL WORLD INTERVALS

Alert Issued Ends 1600 UT 1600 UT	SWI 0001 UT 2400 UT	A <sub>Be</sub> On Days of Alert Period (SWI Underlined)	Number of Flares of IMP ≥ 2 Reported Promptly on Days of Alert Period
1958			
June 03-June 08	June 06-June 08	07-06-06- <u>17</u> - <u>42</u> -08	3-0-4-2-0-3
June 19-June 23	June 20-June 22	10-09- <u>39</u> - <u>27</u> -12	6-1-0-0-1



